

Localized Rapid Ozone Loss in the Northern Winter Stratosphere: An Analysis of UARS Observations

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Observations of low-ozone air pockets forming during northern winter in the middle stratosphere outside the polar vortex provide an opportunity to test models of the photochemistry of ozone at several altitudes, as the trajectories of the associated air parcels are well-defined by Lagrangian transport codes over the periods of interest, and vertical profiles of key species, including ozone, are available from instruments aboard the Upper Atmosphere Research Satellite (UARS). We find that a Lagrangian photochemical model, where the chemistry within an isolated parcel of air is simulated as it travels along a specified trajectory, does reproduce the observed formation of low-ozone pockets in the 6–10 mbar altitude range within the measurement uncertainties, but overestimates the ozone loss rate at higher altitudes. The rapid ozone loss localized in these pockets is due to the isolation of air within anticyclonic regions at high latitudes (low solar zenith angles). Thus the low ozone levels are due to a decrease in the odd oxygen production rate, and not to an increase in the loss rate by reaction with halogen species, as in the “classical” ozone hole.

Abstract.

where transport plays an important role is to employ a Lagrangian photochemical model (Austin *et al.* 1987), where the trajectory of a closed parcel of air (one that does not interact with surrounding air) is explicitly specified using a Lagrangian transport code. In this case, the net change in species abundances in the parcel is solely due to chemistry. The measurements made by the Microwave Limb Sounder (MLS) onboard UARS of the formation of low ozone pockets in the northern winter polar mid-stratosphere (Manney *et al.* 1995) over periods of a few weeks during December 1992 and late February to March 1993 provide an excellent opportunity to apply such a model.

These low ozone air pockets are observed to occur during winter when a strong, long-lived anticyclone forms at high latitudes. In the north, such an anticyclone frequently develops in the Pacific sector (i.e., the "Aleutian high"). As indicated by observations made by MLS, tongues of ozone-rich air are drawn up from low latitudes and into the anticyclone, sometimes having moved around the polar vortex, which in the north is often displaced away from the pole during these periods. These pathways are confirmed by back trajectories computed for air parcels in the low-ozone pockets, which are localized in the 6 to 10 mbar altitude range (Manney *et al.* 1995). As the air lingers at mid-to-high latitudes, its ozone mixing ratio is observed to decrease. The pockets of low ozone lingering in the anticyclone have ozone values as much as 3 ppmv lower than the 8 ppmv ozone mixing ratios in the subtropical regions where the air parcels originated.

Morris *et al.* (1995, 1996) recently investigated the February/March 1993 event with a Lagrangian photochemical model. The conclusion was that air in the low ozone pockets at the 850 K potential temperature level is isolated at high latitudes for periods of time comparable to the photochemical lifetime of ozone. Since the odd oxygen production rate is low at high latitudes, the ozone abundance drops as it relaxes to the photochemical equilibrium value. Air outside the pockets at similar latitudes meanders back and forth between mid and high latitudes over short timescales compared to the

models must, a more direct test of model photochemistry in regions of the atmosphere rather than looking at seasonal variations of spatially averaged ozone fields, as 2-D the model chemistry.

to the influence of transport, which if improperly parameterized, may obscure errors in between such models and observations in this part of the atmosphere, this may be due the polar regions (e.g., Garcia and Solomon 1983). While there is general agreement employed to properly account for the transport of ozone rich air from the tropics to region of the atmosphere. Two dimensional (2-D) photochemical models are typically at lower altitudes (Ko *et al.* 1989), it is difficult to test the model chemistry in this since transport is an important factor in determining the ozone concentration.

surrounding the simulation of chemical processes in current model descriptions. of ClO lower than model predictions. Thus there still appears to be some uncertainties to within 10% from 31 to 38 km altitude (about 10 to 4 mbar), due to measurements chlorine radicals to show that the calculated ozone production and loss rates balance *et al.* (1996) recently used simultaneous measurements of active hydrogen, nitrogen, and production of ozone relative to the observations above about 2 mbar pressure. Jucks *et* Research Satellite (UARS) to compute ozone loss rates and showed an apparent net altitude, Crutzen *et al.* (1995) recently used observations from the Upper Atmosphere

there is a net loss of ozone in the photochemically controlled region above 6 mbar *et al.* 1995, and Crutzen *et al.* 1995). Although most previous efforts concluded that examining this discrepancy (recent work includes Elluszkiewicz and Allen 1993, Siskind This is known as the "ozone deficit" problem, and many studies have been undertaken tend to underpredict the abundance of ozone relative to the observations (WMO 1985).

is in photochemical steady state. Models of this photochemically controlled region between dynamics and chemistry. At altitudes above about 35 km, or 6 mbar, ozone The vertical profile of ozone in the terrestrial stratosphere reflects the interaction

1. Introduction

budgets are constrained by CLAES observations of HNO_3 and ClONO_2 , along with estimates of total NO_y ($= \text{NO} + \text{NO}_2 + \text{NO}_3 + \text{HNO}_3 + 2\text{N}_2\text{O}_5 + \text{ClONO}_2$) and Cl_y ($= \text{Cl} + \text{ClO} + \text{HCl} + \text{ClONO}_2$) made from the observed abundances of tracer species such as N_2O or CH_4 from CLAES (see, for example, Fahey *et al.* 1990, Plumb and Ko 1992, and Woodbridge *et al.* 1994). The following section describes the spacecraft instruments and the measurements.

2. UARS Observations of Trace Species

The MLS instrument measures thermal emission in the microwave region of the spectrum at the atmospheric limb. A summary of the instrument is given by Barath *et al.* (1993). The Microwave Limb Sounder instrument has three radiometers that measure emission from the atmospheric limb at 63, 183, and 205 GHz. The 63 GHz radiometer provides information on pressure and temperature based on O_2 emission, the 183 GHz radiometer measures emission from H_2O and O_3 , and the 205 GHz radiometer measures bands of ClO and O_3 . At 1 to 10 mbar range, the 205 GHz ozone retrievals have an estimated precision and accuracy of 0.2 to 0.3 ppmv, which is better than 10 percent at 1 mbar, and better than 5 percent at 10 mbar. The 183 GHz retrievals are less accurate due to the poorer characterization of the 183 GHz ozone band. We have employed maps made from the 205 GHz measurements (version 3) for this work. Validation of the MLS ozone measurements is discussed more fully in Froidevaux *et al.* (1996).

The validation of the H_2O measurements is described in Lahoz *et al.* (1996). In the 1 to 10 mbar range of interest, the single profile precision is 0.1 to 0.2 ppmv (better than 4%) and the accuracy is 0.5 to 0.7 ppmv, or better than 10%.

Observations of chlorine monoxide (ClO) are also available from MLS (Waters *et al.* 1996). Unfortunately, the data are quite noisy; the typical single profile noise at the 1σ level ranges from 0.4 ppbv at 10 mbar to 1.2 ppbv at 1 mbar. Typical model peak

MLS measurements of water vapor and ozone. The active nitrogen and chlorine stratospheric abundance of odd hydrogen (HO^+) radicals is constrained by simultaneous al. 1986, McElroy and Salawitch 1989, Allen and Delitsky 1991a). In this paper, the taken using data from both LIMS and ATMOS (e.g., Natarajan *et al.* 1986, Callis *et* in the calculations of the ozone profile above 6 mb. Similar approaches have been H_2O , NO_2 , and temperature to constrain the abundances of HO^+ and NO^+ radicals production and loss rates. Froidenvaux *et al.* (1989) also used LIMS observations of data from the Limb Infrared Monitor of the Stratosphere (LIMS) to compute ozone species. Jackman *et al.* (1986) incorporated O_3 , H_2O , HNO_3 , NO_2 , and temperature be determined by either direct measurements or inferred from measurements of other participate in the catalytic cycles which destroy odd oxygen in the stratosphere may The abundances of the active hydrogen, nitrogen, and chlorine radicals that

Phase photochemistry.

observed and predicted ozone loss rates will highlight uncertainties in the standard gas of species involved in the chemical destruction of ozone. Discrepancies between the (CLAES) instruments along these trajectories, to better constrain the abundances measurements taken by the MLS and the Cryogenic Limb Emission Array Spectrometer our initial study (Nair *et al.* 1995). In this computation of loss rates, we have utilized regions with higher ozone. Furthermore, we do this at several different heights, extending leading to the observed low ozone pockets and within parcels leading to neighboring and February/March 1993 events, using a Lagrangian photochemical model to simulate in this paper, we undertake a more comprehensive study of the December 1992 downward and into the polar vortex.

anticyclone is to cause the air to linger at high latitudes, while keeping it from moving relatively high, and the ozone mixing ratio does not fall to low values. The role of the lifetime of odd oxygen. Consequently the average odd oxygen production rate remains

3. Lagrangian Photochemical Model

We have adapted the Caltech/JPL photochemical model (Allen and Delitsky 1991b) to operate in a Lagrangian mode, similar to the work of Austin *et al.* (1987), Jones *et al.* (1990), and Kawa *et al.* (1993). In this mode of calculation, the chemistry of an isolated air parcel with a specified trajectory is simulated, accounting for changes in temperature, pressure, radiation field, and other relevant physical parameters as warranted by the path of travel. Since we have decoupled the chemistry and dynamics, the observed rate of ozone loss is solely due to photochemical processes. In this case, the continuity equation simply becomes

$$\frac{\partial n_i}{\partial t} = P_i - L_i$$

where n is the number density of species i , and P_i and L_i are the chemical production and loss terms, respectively, for species i . The dataset of reactions and photochemical cross sections, partially shown in Table 1, is taken mostly from the 1994 JPL evaluation (DeMore *et al.* 1994).

Table 1

Back trajectory calculations are run using winds from the United Kingdom Meteorological Office (UKMO) troposphere-stratosphere data assimilation system (Swinbank and O'Neill 1994). The validity of these back trajectory calculations was demonstrated by showing that long lived tracers, particularly N₂O as observed by CLAES, are generally conserved along these paths during the periods when the low-ozone pockets develop (Manney *et al.* 1995).

Parcel positions (longitude, latitude, and pressure) are computed at noon GMT at each day over the period of the trajectory. Profiles of ozone at the computed parcel positions are generated for noon GMT from asynoptic maps, as described by Elson and Froidevaux (1993). Temperature profiles at these times are also derived from UKMO data. The radiation field is computed assuming a spherical atmosphere. The solar flux is taken from Mentall *et al.* (1981) and Mount and Rottman (1982).

mixing ratios of ClO for the air parcels we consider here are on the order of 0.4 ppbv or less. Given the large uncertainties in the single profile values of ClO and its short lifetime, we do not assimilate these observations but instead allow the photochemical lifetime, we do not assimilate these observations but instead allow the model to compute ClO.

CLAES also measures emission spectra at the atmospheric limb, and is further described by Rocche *et al.* (1993). Validated data products from CLAES include CH₄, N₂O, HNO₃, and ClONO₂. The quality of CH₄ and N₂O measurements are discussed by Rocche *et al.* (1996). In the 3 to 15 mbar range, CLAES CH₄ appears to be systematically 15 to 25 percent too high relative to correlative measurements. CLAES agrees to within 20 percent with correlative measurements, with no clear systematic bias as with CH₄. Random errors for each gas are on the order of 5 to 10 percent. The validation of the HNO₃ measurements is discussed in Kumer *et al.* (1996). In the current version of the CLAES data (version 7), HNO₃ appears to be underestimated relative to correlative measurements by 15 to 25% at low mixing ratios (below 8 ppbv). For this reason, we have increased the observed HNO₃ mixing ratios by 20% at all altitudes before assimilating them into the model.

The validation of ClONO₂ measurements is discussed in Mergenthaler *et al.* (1996). In the current version of the CLAES data (version 7), ClONO₂ appears to be underestimated by other gases (chiefly O₃ and CO₂) becomes appreciable and the level interference by other gases (chiefly O₃ and CO₂) becomes appreciable and the uncertainty approaches 100% at 6 mbar. We have incorporated the observations of ClONO₂ into the model for trajectories ending at the 740 K and 840 K levels, since these trajectories remain at higher pressures than 6 mbar for the most part. Calculations along trajectories ending at the 960 K potential temperature level do not assimilate ClONO₂ measurements but instead allow the model to determine ClONO₂.

shown in Figure 1. The high NO_y value is more appropriate for tropical air, whereas the low value is more appropriate for midlatitude air (see Figure 6 of Plumb and Ko). Figure 1

Similar correlations between CH_4 and NO_y as well as Cl_y have also been noted (M. Gunson, R. Salawitch, private communication). The predicted NO_y and Cl_y from the CH_4 relations agree fairly well with the predictions from the N_2O relations, once the observed CH_4 mixing ratio is reduced to account for the measurement bias described earlier.

In order to determine the partitioning between members of the active nitrogen and chlorine families at the start of the run, all of the NO_y and Cl_y is assumed initially to be in the form of HNO_3 and HCl , respectively. The abundances of O_2 , N_2 , and O_3 are fixed to their initial values. All other species concentrations are computed self-consistently. The parcel is held at the latitude, pressure, and season at which it originated for three weeks of model time, long enough for the partitioning between members in the active nitrogen and chlorine families to reach a diurnally varying photochemical equilibrium.

Observations of trace gases (O_3 , H_2O , N_2O , HNO_3 , CH_4 , and ClONO_2) made by UARS instruments over the period of interest have been assimilated into the photochemical model. Using the proper ozone profile is important since models of the terrestrial stratosphere have generally underpredicted ozone abundances in the photochemically controlled region above 35 km altitude (Eluszkiewicz and Allen 1993). The chemical destruction rate of ozone is itself dependent upon the ozone concentration. Thus in order to compute the loss rate without introducing potential errors related to the ozone deficit problem, the observed ozone concentration in the parcel should be used. A second order effect is the sensitivity of the radiation field to the ozone profile, which impacts the abundances of active radical species.

Other species observed by UARS are also assimilated into the model. Although assimilating H_2O , N_2O , and CH_4 should not have much of an impact as they are inert tracers, updating HNO_3 and ClONO_2 might be important as they will affect the

The chemical composition of the parcel is initialized to be consistent with UARS observations. The total NO_y and Cl_y budgets are initialized according to piecewise linear fits made to Figures 6 and 7 in Plumb and Ko (1992). We assume that the parcel is in photochemical equilibrium for the initial latitude in the trajectory file. Since there is some scatter in the NO_y relation, we will use both a high and a low NO_y value, as observational fits made to Figures 6 and 7 in Plumb and Ko (1992).

4. Assimilation of UARS Data

reaction (Johnston and Podoliske 1978, Allen et al. 1984).

We will compare the computed changes in ozone with the observed changes. The net production rate of odd oxygen per day is computed by summing the rates of all reactions in the model that either produce or destroy odd oxygen (defined as the sum of O , O_1D , and O_3), as listed in Table 1. ΔO^x is the net change in odd oxygen per

Table 2
section. Concentrations of all other species in the model (Table 2) are computed from

the continuity equation. Mixing ratios of all species at each level are assumed to be also constructed by linear interpolation. Mixing ratios of O_2 , N_2 , O_3 , H_2O , N_2O , HNO_3 , and CH_4 , are constrained to follow the observed quantities, as described in the following account for vertical motion of the parcel. Updated ozone and temperature profiles are recalculated at each time step with the assumption of hydrostatic equilibrium, to

also updated to account for motion in longitude. The vertical gridding of the atmosphere between the positions in the trajectory code output. The local solar time of the parcel is

position of the parcel (longitude, latitude, log pressure) is found by linear interpolation or sunset), where it is taken to be ten minutes. At the beginning of each time step, the

photochemical model time step is one hour except at twilight (within an hour of sunrise

level 20, and the continuity equation for each species is solved only at this level. The

which is the vertical range of the UKMO analyses. The level of the parcel is fixed to be

The full model atmosphere consists of 40 levels from the surface to 0.316 mbars,

track observations, and one is made from only descending track observations. Each latitude on each map thus has one specific local time associated with it. Mixing ratios of ClONO₂ at the parcel location are assimilated into the model twice per day at the local times corresponding to the parcel latitude; at one local time the observation is read off of the map made using only ascending track observations, at the other the observation is read from the map made from descending track observations. At the same time, we update the HCl and ClO concentrations by holding the ClO/HCl ratio constant. Then assuming that the total Cl_y is fixed, we solve for the abundances of HCl and ClO.

5. Simulation of the February/March 1993 Event

5.1. Baseline Model Results

Figure 2 shows two vertical profiles of ozone taken from an asynoptic map generated from MLS observations for noon GMT, March 7, 1993. The dashed line is inside the low ozone pocket at 258.5° east longitude and 67.5° north latitude. The solid line shows the vertical profile of ozone outside the pocket at 178.5° east longitude and 67.5° north latitude. Note that the ozone peak in the pocket is lower and sharper than for the ozone peak outside the pocket and that the ozone in the pocket is considerably depleted relative to the ozone outside from about 1-10 mbar. The nature and history of these low ozone pockets is described more fully in Manney *et al.* (1995).

Figure 2

We have computed the chemical evolution of parcels along several different trajectories ending at the 740, 840, and 960 K potential temperature levels at the position of the pocket (corresponding to 14.7, 10.4, and 7.1 mbar, respectively, on March 7), and also at the 840 K level outside the pocket (11.3 mbar) at the same latitude. We will compare the magnitudes and trends of the observed and modeled ozone loss rates in parcels leading to the low ozone pocket as well as for parcels with trajectories ending outside the low ozone pocket, where the ozone has “normal” values with the observed

for each day over the time period of the trajectory; one is made using only ascending once during the descending track. Two ClONO_2 concentration maps are produced latitude is observed at two local times per day: once during the ascending track, and must be treated differently. Due to the nature of the UARS orbit, any particular Short lived species that have significant diurnal variations (in this case ClONO_2) the diurnal variation of the short lived species once per day.

partitioning is updated. By updating the NO_y family in this manner, we only disturb partitioning is updated. By updating the NO_y family in this manner, we only disturb (al. 1989). The second is that the $\text{NO}_2/\text{N}_2\text{O}_5$ ratio remains constant before and after the Brassieur and Solomon 1986, Natarajan *et al.* 1986, Callis *et al.* 1986, and Froidevaux *et coefficient for the reaction of NO and ClO (similar expressions have been used by where k_1 is the rate coefficient for the reaction of NO and O_3 , and k_2 is the rate*

$$\frac{[\text{NO}_2]}{[\text{NO}]} = \frac{k_1 [\text{O}_3] + k_2 [\text{ClO}]}{k_1}$$

the NO/NO_2 ratio is given by make two assumptions. The first is that NO is in photochemical steady state such that species to keep the NO_y abundance fixed to its initial value. In order to do this, we measured from the beginning of the run), we update the partitioning of the short lived abundances of these short lived species using standard chemistry. Then every 24 hours abundance does not change appreciably over a day, it is important to conserve the total amount of active nitrogen. Over most of the day, we allow the model to compute the abundance of the NO_y family over a day, it is important to conserve the total

Although the HNO_3 concentration varies slowly enough that the total NO_y short lived species are not otherwise allowed to vary.

maps made for noon GMT for each day over the period of the trajectory. The long concentrations are found by linear interpolation in time between concentrations taken parallel are updated at each time step in the same manner as the ozone abundance. The concentrations of long-lived species (H_2O , N_2O , HNO_3 , and CH_4) in the partitioning of the NO_y and Cl_y families.

2.55 ppbv Cl_y, based on the high NO_y curve of Figure 1. Calculations made using the low NO_y correlation will be addressed in a later section. The mixing ratios of O₂, N₂, and O₃ are fixed to their initial values (20.9%, 78.1%, and 9.1 ppmv, respectively). All other species are computed self-consistently for three weeks of model time such that the members of the NO_y and Cl_y families converge to a diurnally repeating photochemical steady state.

Figure 4a shows the average observed ozone abundance along the 840 K trajectories leading to the pocket. The ozone values have been smoothed by replacing each point with the average of the point and its two neighbors, keeping the endpoints unchanged. This makes it easier to see trends in the day-to-day change in the measured ozone. The second panel compares the day-to-day change in the smoothed ozone (triangles) with the model predictions for change in ozone (open circles) over the period of the trajectory. We estimate the minimum uncertainty on the observed day-to-day change in the ozone mixing ratio to be on the order of 0.4 ppmv; since the precision of individual MLS ozone measurements is on the order of 0.2 ppmv. Note that this uncertainty is in most cases larger than the magnitude of the day-to-day change! Nonetheless, there is an obvious long term trend in the observations, and both the observed and model ozone changes are systematically negative and of roughly the same magnitude, with the notable exception of days 50 through 55, where the model ozone loss is greater than the observed loss. The observed overall decrease in ozone is 3.4 ppmv with an estimated uncertainty of 0.9 ppmv, based on the variation of observed ozone mixing ratios in our “uncertainty box”. The integrated model loss is 3.9 ppmv. The third and fourth panels show the pressure and latitude histories of the parcel, respectively. The fourth panel also shows the average noontime solar zenith angle of the parcels as a function of time, which is a measure of the solar insolation received.

Figure 5 is a similar plot for the trajectories leading to the point outside the pocket at 840 K. The ozone mixing ratio remains relatively constant over this time

Figure 4

Figure 5

is approximately 120 ppbv. We have initialized the parcels with 15.6 ppbv NO_x and The mean N_2O mixing ratio in the cluster of parcels ending at 840 K within the pocket longitude are fixed to their values at noon GMT February 14, 1993 for each trajectory.

To initialize the NO_x and Cl_x partitioning, the latitude, pressure, and solar

in mixing ratios of ozone and other species along the trajectory.

in potential temperature. We will employ this "uncertainty box" to estimate the errors we take the uncertainty to be 5 degrees in longitude, 2.5 degrees in latitude, and 30 K in February 14, 1993, which is the end of the back calculation started on March 7, 1993, species. As a conservative estimate of the cumulative uncertainty in parcel position on pressure, temperature) but also in the initialization and assimilation of spatially variable themselves not only by affecting the physical conditions of the parcel (solar zenith angle, Uncertainties in the trajectory model are difficult to quantify. These errors manifest are high.

ending at the position of the pocket begin at lower latitudes, where ozone abundances levels and at 840 K outside the pocket, as shown in Figure 3. Note that the trajectories trajectories ending in a cluster within the pocket at each of the 740 K, 840 K, and 960 K

but from 150 to 199.5 east longitudes. We have simulated the chemistry along five more representative of the zonal mean, was made for the same dates and latitude range, 71.5 degrees north latitude. A second run for air outside the pocket, with ozone values latitude longitude grid, running from 230 to 279.5 degrees east longitude and 56 to was initialized on the 740, 840, and 960 K isentropic surfaces on a 0.5 by 0.5 degree on March 7, 1993, and ending on February 14, 1993. The first group of trajectories trajectory model was run backwards in time for 22 days, starting at 12 noon GMT (1995); trajectories ending outside the pocket are calculated in a similar manner. The The trajectories ending in the low ozone pocket are described by Manney et al. and predicted ozone loss rates will reflect uncertainties in the model chemistry. changes in ozone along the computed trajectories. Disagreement between the observed

the low NO_y mixing ratios and the asterisks show the computed ozone loss with low NO_y and a modified value for the reaction of OH and ClO. Calculations with modified chemistry will be addressed in a later section.

The results we have presented thus far do show a qualitatively interesting point, namely that the model ozone deficit still appears to exist above about 7 mbar altitude, while the model appears to predict changes in ozone below this altitude reasonably accurately. However, as we have seen for all of the cases shown above, there are many uncertainties. For example, individual trajectories may not properly describe the true air motion, particularly near the end of the back trajectory (day 45), or the initialization of species in the parcel may be wrong.

Figure 9 compares the simulated daily changes in ozone with the observed daily changes in each group of trajectories. Each group has (5 trajectories)*(22 points/trajectory) = 110 points in it. We have plotted the average of the day-to-day changes for each group along with the sample standard deviation. Note that the predicted ozone decreases for parcels ending at 960 K are all much greater than the observed losses. It must be noted that this is a relatively small sample size and more cases should be run before quantitative conclusions may be drawn. However, it does appear that the model adequately describes the chemistry in the parcels below about the 840 K potential temperature level and consistently overpredicts ozone loss rates above this.

Figure 9

5.2. Influence of Vertical and Latitudinal Motion on Ozone Loss Rates

A trajectory has three components, motion in longitude, motion in latitude, and motion in pressure. Since the motion in longitude has a small effect on the solar zenith angle, compared to the rotation of the earth, motion in latitude and pressure have a much larger impact on changing the radiation field, which in turn drives the processes that produce and destroy ozone. We have seen from the 960 K runs in the previous

We show three sets of computed ozone losses; the dark circles are computed using where there is an ozone deficit to one where there is not, reflected in the second panel. models at altitudes above about 6 mbar. This parcel appears to move from an altitude This is interesting because the ozone deficit generally manifests itself in photochemical agreement near the end of the trajectory, as the parcel descends to around 7 mbar. calculations. The model predictions for the day-to-day changes come into much better observed ozone mixing ratio, which is fixed to match the observed quantity in the model. This unphysical result occurs because the computed loss rates are based upon the integrated model loss of ozone is 13.9 ppmy, greater than the initial abundance; ozone losses are consistently much greater than the observations would indicate. In fact, at the 960 K potential temperature level. Note in the second panel that the model Finally, Figure 8 summarizes the history of the parcel ending within the pocket change (0.3 ppmy).

ozone are small (~ 0.1 ppmy) compared to the uncertainty on the observed day-to-day losses are small (~ 0.1 ppmy) compared to the uncertainty on the observed day-to-day last day. However, it must still be noted that both the model and observed changes in before day 55. Over the last ten days or so, the model is in better agreement up to the particular trajectory probably does not describe the true motion of this parcel, at least we note considerable variability over the first ten days of the trajectory. Thus this period. Looking at the evolution of the long lived tracers for this parcel (Figure 7), trajectory. This leads us to suspect that the trajectory may be in error for this time the pocket. Note the large drop in the observed ozone over the first few days of the Figure 6 shows the model results at the 740 K potential temperature level within significantly smaller than for the trajectories leading to the pocket.

agreement between model and the observations is poor for this case, both values are of 0.4 ± 0.5 ppmy, while the model predicts a decrease of 1.5 ppmy. Although the observed change in ozone over the three week period of the trajectory is a slight increase period, although a gentle oscillation with a period of about a week can be seen. The

Figures 4 and 5, the parcels follow similar pressure paths, but very different latitude paths, although the final latitude is 67.5° north in each case. The noontime solar zenith angle is greater than 60° for the last twelve days or so for the parcel ending within the pocket (Fig. 4d), while the noontime solar zenith angle mostly remains below 60° for the parcel ending outside the pocket (Fig. 5d). In each case, the production of odd oxygen is anticorrelated with the noontime solar zenith angle (which decreases as the parcel nears the pole). For the parcel ending outside the pocket, there is a slight net production of odd oxygen at low latitudes as the ozone concentration in the parcel is below the low latitude equilibrium value. Conversely, at high latitudes, there is a net loss of ozone. The ozone mixing ratio in parcels rapidly moving between these two regions reflects a balance between the average production and loss over periods comparable to the photochemical lifetime of ozone, which is about 10 days at 35 km altitude (Ko *et al.* 1989). Panel (e) of Fig. 10 shows the calculated ozone change for a parcel following the trajectory ending at the 840 K potential temperature level outside the low ozone pocket, with the same chemical initialization as parcels (a)-(d). The calculated ozone loss is -1.89 ppmv, much smaller than the loss along the trajectory leading to the pocket. Thus the varying concentration of ozone in air parcels at the latitudes of the low-ozone pockets reflects different exposure times to sunlight as the parcels take different paths through the higher latitudes.

Morris *et al.* (1996) performed numerical experiments to examine the effects of varying the trajectory followed by a parcel on its ozone content. Their findings are consistent with those presented here; that the low ozone pockets form in the middle stratosphere when air is held at high latitudes for a week or more, and that air outside the anticyclone does not exhibit these low ozone values as it is not confined to high latitudes. Adjusting other factors, such as temperature and the initial values of ozone and NO_y, in their model did not have an appreciable effect on the final ozone mixing ratios.

We have performed a series of numerical experiments where we examine the effects on the ozone abundance of individually varying the pressure and latitude over one of the trajectories shown in the last section ending at the 840 K level within the pocket. In each model run, we either allow the pressure and/or latitude to vary according to the trajectories shown in the last section ending at the 840 K level within the pocket. We have performed a series of numerical experiments where we examine the effects in pressure and latitude on the loss rate of ozone.

In each model run, we either allow the pressure and/or latitude to vary according to the computed trajectory (shown in Figures 4c and 4d), or fix them to their initial values (6.74 mbar and 22.9° north, respectively). In these calculations we allow the model to compute the abundances of all species, including ozone, self-consistently. Since these are fictitious trajectories, it is not appropriate to assimilate measurements from the actual trajectory. The chemical initialization for each parcel is the same as for the 840 K trajectory in the previous section. The computed ozone mixing ratios and total change in ozone for each case are shown in Figure 10.

We see that allowing the pressure to vary over the range traversed by the parcel compares well with the observed change of -3.4 ppmv.

Note, incidentally, that the calculated ozone change in case (d) is -3.6 ppmv, which poleward of 60° for most of the last ten days, the equilibrium value is around 5 ppmv. In cases (c) and (d), where the parcel rapidly moves northward and stays 8 ppmv. In cases (a) and (b), where the latitude is fixed at 22.9°, the equilibrium value of ozone is about 8 ppmv. In case (a) due to the lower solar insolation and hence lower oxygen photolysis rate. In cases (a) and (b), where the latitude is fixed at 22.9°, the equilibrium value of ozone is about 8 ppmv. This photochemical equilibrium value is lower closer to the pole, production equals loss. The ratio of ozone in the parcel relaxes towards a photochemical equilibrium value, where comparable to the observed change along the actual trajectory. In each case, the mixing ratio of ozone in the parcel is held fixed. However, allowing the latitude to vary produces a large change in ozone, produces a negligible change in the ozone mixing ratio from the cases where the pressure is held fixed. We see that allowing the pressure to vary over the range traversed by the parcel due to the lower solar insolation and hence lower oxygen photolysis rate. In cases (a)

(Figure 11). After this, the parcel remains at high latitudes where zonal means are not particularly meaningful due to the pronounced zonal asymmetry.

Figure 11

Figure 8b shows the net change (production minus loss, or $P - L$) in ozone over the parcel trajectories ending at 960 K for the standard case, low NO_y , and low NO_y plus the modified $\text{OH} + \text{ClO}$ reaction. Decreasing the NO_y and ClO contribute roughly equally to reducing the net ozone loss at high altitudes. Note that although the combination of the low NO_y and 7% HCl branch causes a considerable decrease in $P - L$ relative to the standard case, particularly in the early part of the trajectory when the parcel is higher than the 5 mbar altitude level, the integrated ozone loss over the period of the trajectory is 10.1 ppmv, still greatly in excess of the observed 2.56 ppmv. Thus we find only slight improvement in the agreement between modeled and observed ozone loss rates, as shown in Figure 9. There still appears to be a clear discrepancy between model and observed ozone chemistry in the upper stratosphere.

Jucks *et al.* (1996) used simultaneous measurements taken in September 1989 at 34° north of active hydrogen, nitrogen, and chlorine radicals to show that the calculated ozone production and loss rates balance to within 10% from the 10 to 4 mbar altitude levels. At the present time we have no explanation for the discrepancy between their findings and ours, namely that model error in the Cl_y partitioning does not appear to be entirely responsible for the 960 K ozone deficit.

6. Simulation of the December 1992 Event

A low ozone pocket was also seen in December 1992, however, the latitude of formation was not as near the pole; the center of the pocket was located at about 45° north as opposed to about 65° north latitude for the February/March 1993 event. Figure 12 shows observed vertical profiles of ozone inside the pocket (174.5° east, 47.5° north) and outside the pocket (229.5° east, 47.5° north) on December 24, 1992. Again, the ozone at the position of the pocket has a low, sharp peak, and is considerably

Figure 12

OH reaction up to about day 54, where the parcel is located between 4 and 5 mbar longitude and pressure) is greatly improved after the incorporation of the new ClO + ClO and the zonal mean value from MLS observations (interpolated to the appropriate branch at the initial parcel latitude. The agreement between the model daily maximum chlorine species was found by running the model to steady state with the 7% HCl run is assumed to be the same as in the high NO_y case. The initial partitioning between the 7% HCl branch. The relative partitioning between NO_y species at the start of the We have made model runs at the 960 K level with the low NO_y distribution and additional HCl source.

partitioning. The work of Jucks *et al.* (1996) also strongly supports the existence of an ClO and OH markedly increases the agreement between the modeled and observed Cl_y (1996) found that incorporation of a 7% production yield of HCl from the reaction of from February to March of 1993 between 40 and 46 km altitude. Michelsen *et al.* showed that model ClO is a factor of two higher than MLS measurements obtained which results in a greater degree of ozone loss in the models. Dessler *et al.* (1996) ClO/HCl ratio is overpredicted by as much as a factor of three near 40 km altitude, photochemical models does not agree with observations below 50 km altitude. The active chlorine; Michelsen *et al.* (1996) showed that the partitioning of Cl_y in current Another possible candidate for the high model ozone losses is an overabundance of consider cases where the parcels are initialized with the lower NO_y distribution.

appropriate for parcels originating from the tropics (equatorward of 30°). We will now have employed the higher NO_y distribution (solid line of Figure 1, top panel), which is particularly at N₂O values around 50 ppbv. All of the model runs presented thus far correlation we use is based on Figure 6 of Plumb and Ko (1992), which has some scatter, the NO_y in the parcel may be overestimated. As mentioned earlier, the N₂O - NO_y One reason for the enhanced model loss early in the 960 K trajectories may be that

5.3. Uncertainties in NO_y and Cl_y.

much larger for the trajectories leading to the pocket. The parcels ending outside the pocket are subjected to higher solar zenith angles over the last three days of the model simulation but there is no decrease in ozone since this is a short period of time compared to the photochemical lifetime of ozone. Thus it appears that the same processes which produce a low ozone pocket at high latitudes, also operate at lower latitudes.

The 960 K results are summarized in Figure 17. We find that the model prediction of 6.1 ppmv for the integrated ozone loss exceeds the average observed losses of 1.9 ± 0.4 ppmv for parcels ending at the 960 K potential temperature level, although the discrepancy is not as great as for the February/March 1993 case. This is because the ozone mixing ratios in the parcel over the period of the trajectory were lower than for the February/March 1993 event (Manney *et al.* 1995), which leads to smaller loss rates. Even so, there remains a serious discrepancy between observed and modeled loss rates, particularly in the early part of the trajectories where the parcels are above about 6 mbar.

Fig. 17b also shows the computed daily ozone losses with low NO_y and low NO_y plus the modified OH + ClO reaction. The integrated losses in these cases were -5.1 and -4.5 ppmv, respectively. Again, although the agreement between the observed and modeled ozone losses is improved, the model still overestimates the ozone loss, particularly at lower pressures early in the trajectory. In this instance, the model also overestimates the ClO mixing ratio early in the trajectory, compared to the zonal mean ClO measurements from MLS (Figure 18), unlike the good agreement found in the other event. Thus even with the 7% branch for HCl production, the description of chlorine chemistry in the model is still in error. If we extrapolate the effect of lowering ClO on the integrated ozone loss, we estimate that lowering ClO by another 20% to agree with the MLS zonal means would equivalently reduce the magnitude of the computed ozone loss by 20%, or to about 3.5 ppmv, still leading to a discrepancy between observed and computed ozone losses.

Figure 17

Figure 18

Figure 16

solar zenith angles were actually slightly greater since this event occurred close to the northern winter solstice (see Figs. 14-17). Figures 14 and 15 show the average observed ozone mixing ratio, observed and calculated loss rates, pressure, latitude, and noon solar zenith angle for the parcels ending at the 740 and 840 K potential temperature levels, respectively, for the December 1992 event. The computed model loss of -1.4 ppmy is smaller in magnitude than the observed -2.7 ± 1.0 ppmy for the 740 K trajectory. The integrated model ozone loss of 2.8 ppmy is in good agreement with the observed ozone loss of -2.7 ± 0.6 ppmy for the

Figure 15

As with the February/March 1993 event, the ozone loss rates along trajectories depleted relative to the extratropical profile at the 10 mbar level. Leading to both inside and outside the pocket were computed. Figure 13 shows the calculated trajectories for five parcels ending at each of the 740, 840, and 960 K leading to both inside and outside the pocket were computed. Figure 13 shows the potential temperature levels within the pocket (corresponding to 15.0, 10.6, and 7.0 mbar, respectively, on December 24), and for five parcels ending at the 840 K potential temperature level outside the pocket (10.6 mbar). The trajectories for this event differ from the February/March 1993 event since the air parcels leading to the pocket stay close to 45° north latitude for most of the period of the trajectory. Although the latitude of this event was about 20° lower than for the February/March 1993 event, the noontime solar zenith angles were actually slightly greater since this event occurred close to the

Figure 13

ratios found in the vortex and the low ozone pockets. The low ozone levels for parcels outside the polar vortex are due to the longer residence times of these parcels at higher latitudes and the corresponding decrease in the odd oxygen production rate, and not to an increase in the loss rate by reaction with halogen or other species. These findings are consistent with the conclusions of Morris *et al.* (1996).

At altitudes below about 7 mbar, the photochemical model satisfactorily reproduces the observed ozone loss. However, we find that the photochemical model consistently overestimates the loss rate for ozone at pressures less than about 7 mbar. This is a region of the atmosphere where models have historically overpredicted ozone loss rates relative to production (the "ozone deficit"). In contrast to earlier work, we find that adjusting the chlorine partitioning does not definitively resolve this issue above the 960 K potential temperature level. Thus it appears that significant uncertainties remain in our understanding of the photochemical processes controlling ozone in the upper stratosphere.

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Figure 19

Figure 19 summarizes the comparison between the observed daily changes in ozone with the model changes similar to Figure 9. Like the February/March 1993 event, we find that the model agrees well with the observed loss rates at the lower altitudes, but the model overestimates the loss rate at the highest altitude. Incorporation of lower NO_x abundances and the 7% production yield of HCl from ClO + OH does lead to better agreement between the model and observed ozone loss rates, although this agreement is at the extreme ranges of the uncertainties.

We have investigated the chemical evolution of specific parcels of air along trajectories computed by Manney *et al.* (1995) leading to observed pockets of

extra-vertically, ozone poor air in the middle stratosphere in February/March 1993 and December 1992. These pockets of air occur in an altitude region where both dynamics and chemistry are important contributors to the ozone budget. Observations of trace species from the Microwave Limb Sounder (MLS) and the Cryogenic Limb Array Etalon Spectrometer (CLAES) aboard the Upper Atmosphere Research Satellite (UARS) were incorporated into a Lagrangian photochemical model to test whether the description of

photochemical processes in current models accurately determines the rate of ozone loss near the 740, 840, and 960 K potential temperature levels. Trajectory calculations for both events indicate that when air is held at high solar zenith angles for periods of time longer than the photochemical lifetime of ozone, a low ozone pocket forms. Since the photolysis rate of oxygen at high solar zenith angles is small, ozone mixing ratio falls to low values in order for the loss rate to balance the production of odd oxygen. Air outside of these pockets and outside the polar vortex typically swings back and forth between mid and high latitudes such that the average

production rate of odd oxygen is high relative to that of air in the low ozone pocket. Thus the steady state ozone mixing ratio for this "normal" air is higher than the mixing production rate of odd oxygen is high relative to that of air in the low ozone pocket. Trajectory calculations for both events indicate that when air is held at high solar zenith angles for periods of time longer than the photochemical lifetime of ozone, a low ozone pocket forms. Since the photolysis rate of oxygen at high solar zenith angles is small, ozone mixing ratio falls to low values in order for the loss rate to balance the production of odd oxygen. Air outside of these pockets and outside the polar vortex typically swings back and forth between mid and high latitudes such that the average

7. Conclusions

The model overestimates the loss rate at the highest altitude. Incorporation of lower NO_x abundances and the 7% production yield of HCl from ClO + OH does lead to better agreement between the model and observed ozone loss rates, although this agreement is at the extreme ranges of the uncertainties.

Figure 19 summarizes the comparison between the observed daily changes in ozone

- 1009–1032, 1983.
- Crutzen, P.J., J.-U. Groß, C. Brühl, R. Müller, and J.M. Russell III, A reevaluation of the ozone budget with HALOE UARS data: No evidence for the ozone deficit, *Science*, **268**, 705–708, 1995.
- Demore, W.B., S.P. Sander, D.M. Golden, R.F. Hampson, M.J. Kurylo, C.J. Howard, A.R. Ravishankara, C.E. Kolb, and M.J. Molina, Chemical Kinetics and Photochemical Data for Use in Stratospheric Modeling, Evaluation Number 11 *JPL Publication 94-26*, 1994.
- Dessler, A.E., S.R. Kawa, D.B. Considine, J.W. Waters, L. Froidevaux, and J.B. Kumer, UARS measurements of ClO and NO₂ at 40 and 46 km and implications for the model “ozone deficit”, *Geophys. Res. Lett.*, **23**, 339–342, 1996.
- Elson, L.S., and L. Froidevaux, The use of Fourier transforms for asynoptic mapping: Early results for the Upper Atmosphere Research Satellite Microwave Limb Sounder, *J. Geophys. Res.*, **98**, 23039–23049, 1993.
- Eluszkiewicz, J.E. and M. Allen, A global analysis of the ozone deficit in the upper stratosphere and lower mesosphere, *J. Geophys. Res.*, **98**, 1069–1082, 1993.
- Fahey, D.W., S. Solomon, S.R. Kawa, M. Loewenstein, J.R. Podolske, S.E. Strahan, and K.R. Chan, A diagnostic for denitrification in the winter polar stratosphere, *Nature*, **345**, 698–702, 1990.
- Froidevaux, L., M. Allen, S. Berman, and A. Daughton, The mean ozone profile and its temperature sensitivity in the upper stratosphere and lower mesosphere: An analysis of LIMS observations, *J. Geophys. Res.*, **94**, 6389–6417, 1989.
- Froidevaux, L., W.G. Read, T.A. Lungu, R.E. Cofield, E.F. Fishbein, D.A. Flower, R.F. Jarnot, B.P. Ridenoure, Z. Shippony, J.W. Waters, J.J. Margitan, I.S. McDermid, and R.A. Stachnik, Validation of UARS MLS Ozone Measurements, *J. Geophys. Res.*, **101**, 10017–10060, 1996.
- Garcia, R.R., and S. Solomon, A numerical model of the zonally averaged dynamical and chemical structure of the middle atmosphere, *J. Geophys. Res.*, **88**, 1379–1400, 1983.
- Jackman, C.H., R.S. Stolarski, and J.A. Kaye, Two-dimensional monthly average ozone balance from Limb Infrared Monitor of the Stratosphere and Stratospheric and Mesospheric

References

- Allan, M., J.I. Lunine, and Y.L. Yung, The vertical distribution of ozone in the mesosphere and lower thermosphere, *J. Geophys. Res.*, 89, 4841-4872, 1984.
- Allan, M., and M.L. Delitsky, Imaging the abundances of ClO and HO₂ from Spacelab 3 Atmospheric Trace Molecule Spectroscopy, A test of odd-oxygen photochemistry using Spacelab 3 comparisons with LIMS observations in a stratospheric trajectory coordinate system, *Q. J. R. Meteorol. Soc.*, 113, 361-392, 1987.
- Austin, J., R.C. Palister, J.A. Pyle, A.F. Tuck, A.M. Zavodny, Photochemical model 12883-12891, 1991b.
- Allen, M., and M.L. Delitsky, Atmospheric Trace Molecule Spectroscopy observations, *J. Geophys. Res.*, 96, 2913-2919, 1991a.
- Allen, M., and M.L. Delitsky, Imaging the abundances of ClO and HO₂ from Spacelab 3 Atmospheric Trace Molecule Spectroscopy observations, *J. Geophys. Res.*, 96,
- Austin, J., R.C. Palister, J.A. Pyle, A.F. Tuck, A.M. Zavodny, Photochemical model 12883-12891, 1991b.
- Bartel, F.T., M.C. Chavez, R.E. Cohen, D.A. Flower, M.A. Frekringe, M.B. Gram, W.M. Harris, J.R. Holden, R.F. Jarrot, W.G. Kloezeeman, G.J. Klose, G.K. Lau, M.S. Loo, B.J. Maddison, R.J. Matrauch, R.P. McKinnell, G.B. Peckham, H.M. Pickett, G. Siebes, F.S. Solis, R.A. Sutte, J.A. Tarasla, J.W. Waters, and W.J. Wilson, The Upper Atmosphere Research Satellite Microwave Limb Sounder Instrument, *J. Geophys. Res.*,
- Brasseur, G., and S. Solomon Aeronomics of the middle atmosphere, second edition, D. Reidel, 98, 10751-10762, 1993.
- Calles, I.B., M. Natarajan, R.E. Boughez, J.M. Russell III, and J.D. Lambeth, Stratospheric distributions, *J. Geophys. Res.*, 91, 1167-1197, 1986.
- Chamce, K., W.A. Traub, D.G. Johnson, K.W. Jucks, P. Ciarpalli, R.A. Stachnik, R.J. Salawitch, and H.A. Michelsen, Simultaneous measurements of stratospheric HO_x, NO_x, and Cl_x: Comparison with a photochemical model, *J. Geophys. Res.*, 101, 9031-9043,
- Cruzen, P.J., and U. Schmidli, Chemical budgets of the stratosphere, *Planet. Sp. Sci.*, 31, 1996.

13939–13950, 1995.

McElroy, M.B., and R.J. Salawitch, Stratospheric ozone: Impact of human activity, *Planet. Sp. Sci.*, 37, 1653–1672, 1989.

Mentall, J.E., J.E. Frederick, J.R. Herman, The solar irradiance from 200 to 330 nm, *J. Geophys. Res.*, 86, 9881–9884, 1981.

Mergenthaler, J.L., J.B. Kumer, A.E. Roche, R.W. Nightingale, J.F. Potter, J.C. Gille, S.T. Massie, P.L. Bailey, D. Edwards, P.S. Connell, D.E. Kinnison, M.R. Gunson, M.C. Abrams, G.C. Toon, B. Sen, J.-F. Blavier, D.G. Murcray, F.J. Murcray, and A. Goldman, Validation of CLAES ClONO₂ measurements, *J. Geophys. Res.*, 101, 9603–9620, 1996.

Michelsen, H.A., R.J. Salawitch, M.R. Gunson, C. Aellig, N. Kaempfer, M.M. Abbas, M.C. Abrams, T.L. Brown, A.Y. Chang, A. Goldman, F.W. Irion, M.J. Newchurch, C.P. Rinsland, G.P. Stiller, and R. Zander, Stratospheric chlorine partitioning: Constraints from shuttle-borne measurements of HCl, ClNO₃, and ClO, *Geophys. Res. Lett.*, 23, 2361–2364, 1995.

Morris, G.A., S.R. Kawa, A.R. Douglass, M.R. Schoeberl, J. Waters, and L. Froidevaux, Mid-stratospheric ozone loss in wave breaking events, *Eos, Transactions, American Geophysical Union*, 76, S78, 1995.

Morris, G.A., S.R. Kawa, A.R. Douglass, M.R. Schoeberl, L. Froidevaux, and J. Waters, Low ozone “pockets” explained, Submitted to *J. Geophys. Res.*, 1996

Mount, G.H., and G.J. Rottman, The solar absolute spectral irradiance 1150–3173 Å: 17 May 1982, *J. Geophys. Res.*, 88, 5403–5410, 1982.

Nair, H., M. Allen, G.L. Manney, R. Zurek, and L. Froidevaux, An investigation of low ozone anomalies in the winter middle stratosphere using a Lagrangian photochemical model, *Eos, Transactions, American Geophysical Union*, 76, F80, 1995.

Natarajan, M., L.B. Callis, R.E. Boughner, and J.M. Russell III, Stratospheric photochemical studies using Nimbus 7 data: 1. ozone photochemistry, *J. Geophys. Res.*, 91, 1153–1166, 1986.

Plumb, R.A., and M.K.W. Ko, Interrelationships between mixing ratios of long lived

- pockets in the middle stratospheric anticyclone during winter, *J. Geophys. Res.*, 100,
 Mergenthaler, A.E., Roche, A., O'Neill, and R. Swinbank, Formation of low ozone
 Manney, G.L., L. Froidevaux, J.W. Waters, R.W. Zurek, J.C. Gillie, J.B. Kumer, J.L.
 UARS MLS 183 GHz H₂O Measurements, *J. Geophys. Res.*, 101, 10129-10149, 1996.
 H.C. Pumphrey, W.G. Read, Z. Shippony, R.A. Suttie, J.W. Waters, Validation of
 Lahoz, W.A., M.R. Suttie, L. Froidevaux, R.S. Harwood, C.L. Lau, T.A. Lungu, G.E. Peckham,
 9621-9656, 1996.
- on the NASA Upper Atmosphere Research Satellite (UARS), *J. Geophys. Res.*, 101,
 from the Cryogenic Limb Array Experiment (CLAES) instrument deployed
 and D.G. Johnson, Comparison of correlative data with nitric acid data version v0007
 C.R. Webster, D.G. Murcay, F.H. Murcay, A. Goldmann, W.A. Traub, K.W. Jucks,
 Baile, M.R. Guinason, M.C. Abrams, G.C. Toon, B. Sen, J.-F. Blavier, R.A. Stachnik,
 Kumer, J.B., J.L. Mergenthaler, A.E. Roche, R.W. Nighswander, J.C. Gillie, S.T. Massie, P.L.
 two-dimensional models, *J. Geophys. Res.*, 94, 9889-9896, 1989.
- in determining the stratospheric concentration of ozone in one-dimensional and
 Ko, M.K.W., N.D. Sze, and D.K. Weisenstein, The roles of dynamical and chemical processes
 chemistry along trajectories, *Geophys. Res. Lett.*, 20, 2507-2510, 1993.
- Loewenstein, Interpretation of NO_x/NO_y observations from AASE-II using a model of
 Woodbridge, H. Jonsson, L.R. Latit, P.A. Newmann, M.H. Proffitt, D.E. Anderson, M.
 Kawamura, S.R., D.W. Fahey, J.C. Wilson, M.R. Schoeberl, A.R. Douglass, R.S. Stolarski, E.L.
 Science, , 1996.
- Ozone production and loss rate measurements in the middle stratosphere, Submitted to
 Jucks, K.W., D.G. Johnson, K.V. Chanec, W.A. Traub, R.J. Salawitch, and R.A. Stachnik,
Res. Lett., 17, 545-548, 1990.
- cloud formation on chemical composition during the 1988/89 Arctic winter, *Geophys.*
 Jones, R.L., D.S. McKenna, L.R. Pool, and S. Solomon, On the influence of polar stratospheric
Phys. Space Phys., 16, 491-519, 1978.
- Johnston, H.S., and J. Podolske, Interpretations of stratospheric chemistry, *Rev. Geo-*
 Sounder data, *J. Geophys. Res.*, 91, 1103-1116, 1986.

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- stratospheric constituents, J. *Geophys. Res.*, 97, 10145-10156, 1992.
 Reber, C.A., C.E. Treverathan, R.J. McNeal, and M.R. Luthe, The Upper Atmosphere Research Satellite (UARS) Mission, J. *Geophys. Res.*, 98, 10643-10647, 1993.
 Rocche, A.E., J.B. Kumer, J.L. Mergenthaler, G.A. Ely, W.G. Uhlinger, J.F. Pottier, T.C. James, and L.W. Sterritt, The Cryogenic Limb Array Etalon Spectrometer (CLAES) on UARS: Experiment Description and Performance, J. *Geophys. Res.*, 98, 10763-10775,
 Rocche, A.E., J.B. Kumer, R.N. Nighthingale, J.L. Mergenthaler, P.L. Baile, S.T. Massie, J.C. Gillie, M.R. Gunson, M.C. Abrams, G.C. Toon, C.R. Webster, W.A. Traub, K.W. Jucks, D.G. Johnson, R. de Zafra, D.G. Murcay, F.H. Murcay, and A. Goldman, Validation of CH₄ and N₂O measurements by the CLAES instrument on the Upper Atmosphere Research Satellite, J. *Geophys. Res.*, 101, 9679-9710, 1996.
 Siskind, D.E., B.J. Connor, R.S. Eckman, E.B. Remsberg, J.J. Tsou, and A. Parfis, An intercomparison of model ozone deficits in the upper stratosphere and mesosphere from two data sets, J. *Geophys. Res.*, 100, 11191-11201, 1995.
 Stachnik, R.A., J.C. Hardy, J.A. Tarasla, J.W. Waters, and N.R. Erickson, Submillimeterwave heterodyne measurements of stratospheric ClO, HCl, O₃, and HO₂: First results, J. *Geophys. Res.*, 19, 1931-1934, 1992.
 Swinbank, R., and A. O'Neill, A stratosphere-troposphere data assimilation system, Mon. *Weather Rev.*, 122, 686-702, 1994.
 Waters, J.W., W.G. Read, L. Froidevaux, T.A. Lungeu, V.S. Perun, R.A. Stachnik, R.F. Jarrot, R.E. Coffield, E.F. Fishbein, D.A. Flower, J.R. Burke, J.C. Hardy, L.T. Nakamura, B.P. Ridenour, Z. Shippony, R.P. Thurstans, Validation of UARS MLS ClO Measurements, J. *Geophys. Res.*, 101, 10091-10127, 1996.
 Webster, R.D., May, J.M. Gilligan, S.A. Montzka, K.A. Boering, and R.J. Salawitch, Estimates of total organic and inorganic chlorine in the lower stratosphere from in situ flask measurements during AASE II, J. *Geophys. Res.*, 100, 3057-3064, 1995.

Figure 10. Computed ozone mixing ratios for five cases: (a) pressure = 6.74 mbar, latitude = 22.9° , (b) pressure varies according to Fig. 4c, latitude = 22.9° , (c) pressure = 6.74 mbar, latitude varies according to Fig. 4d, (d) vary pressure and latitude according to Figs. 4c and 4d, (e) vary pressure and latitude according to Figs. 5c and 5d.

Figure 11. Model daily maximum ClO mixing ratios using standard chemistry (solid line) and modified chlorine chemistry (dashed line) for a 960 K parcel. The symbols represent MLS zonal mean ClO interpolated to the latitude and pressure of the parcel, along with the estimated 15% measurement uncertainty. Refer to Figure 8c for the latitude and pressure history of the parcel. MLS zonal means northward of 50° are not shown due to lack of zonal symmetry.

Figure 12. Ozone profiles inside (dashed line) and outside the pocket (solid line) on December 24, 1992. Both profiles are at 47.5° north latitude.

Figure 13. Four sets of trajectories for the December 1992 event. Parcels begin on December 3, 1992 and end on December 24, 1992.

Figure 14. As Fig. 4, but for the trajectories ending at 740 K inside the pocket during the December 1992 event. The observed change in ozone is -2.7 ± 1.0 ppmv and the model change is -1.4 ppmv.

Figure 15. As Fig. 14, but for the trajectories ending at 840 K inside the pocket. The observed change in ozone is -2.7 ± 0.6 ppmv and the model change is -2.8 ppmv.

Figure 16. As Fig. 14, but for the trajectories ending at 840 K outside the pocket. The observed change in ozone is 0.7 ± 0.4 ppmv and the model change is -1.7 ppmv.

Figure 17. As Fig. 14, but for the trajectories ending at 960 K within the pocket. The observed change in ozone is -1.9 ± 0.4 ppmv and the model change is -6.1 ppmv. Computations with modified chemistry are also shown, in the same fashion as Fig. 8.

sample standard deviation in each direction. Potential temperature levels for ease of visualisation. The error bars extend out to one period of the trajectories. Points are slightly offset from the 740, 840, and 960 K Figure 9. Average model (circles) vs. observed (triangles) daily changes in ozone over modified chlorine chemistry (asterisks)

loss with standard chemistry (open circles), low NO_x (dark circles) and low NO_x plus change in ozone is -2.6 ± 0.6 ppbv and the model change is -12.9 ppbv. (b) shows model Figure 8. As Fig. 4, but for the parcels ending at 960 K within the pocket. The observed 740 K trajectories.

Figure 7. Trends in O₃ and the long lived tracer species CH₄, N₂O, and H₂O along the change in ozone is -3.2 ± 1.8 ppbv and the model change is -1.7 ppbv.

Figure 6. As Fig. 4, but for the parcels ending at 740 K within the pocket. The observed change in ozone is 0.4 ± 0.5 ppbv and the model change is -1.5 ppbv.

Figure 5. As Fig. 4, but for the trajectories ending outside the pocket. The observed angle (dashed line) of the parcels as a function of time.

chemistry (open circles). (c) Pressure, (d) latitude (solid line) and noontime solar zenith by MLS. (b) Observed change in ozone (triangles), model calculation with standard is 3.4 ± 0.9 ppbv and the model loss is 3.9 ppbv. (a) Ozone mixing ratio as observed Figure 4. Average ozone loss over 840 K anomalous trajectories. The observed loss by 0.5 degrees. Parcels begin on February 14, 1993 and end on March 7, 1993.

contains five distinct parcels, ending in a cross pattern with the parcel centres separated Figure 3. Four sets of trajectories for the February/March 1993 event. Each group March 7, 1993. Both profiles are at 67.5° north latitude.

Figure 2. Ozone profiles inside (dashed line) and outside the pocket (solid line) on Cl_y.

line) and low (dashed line) total NO_x as a function of N₂O. The lower panel shows total Figure 1. Total NO_x and Cl_y as a function of N₂O. The top panel shows high (solid

Table 1. Important reactions for the odd oxygen balance

	Reaction	ΔO_x	Reaction	ΔO_x	Reaction	ΔO_x	Reaction	ΔO_x
R1	$O_2 + h\nu \rightarrow 2O$	+2						+1
R2	$O_2 + h\nu \rightarrow O + O(^1D)$	+2	R147	$N + NO \rightarrow N_2 + O$				+1
R5	$H_2O_2 + h\nu \rightarrow OH + O$	+1	R148	$N + NO_2 \rightarrow N_2O + O$				-1
R8	$H_2O + h\nu \rightarrow 2H + O$	+1	R149	$NO + O + M \rightarrow N_2O_2 + M$				-1
R10	$NO + h\nu \rightarrow N + O$	+1	R150	$NO + O_3 \rightarrow N_2O_2 + O_2$				-1
R11	$NO_2 + h\nu \rightarrow NO + O$	+1	R158	$NO_2 + O_3 \rightarrow N_2O_3 + O_2$				-1
R12	$NO_3 + h\nu \rightarrow NO_2 + O$	+1	R166	$NO_3 + O \rightarrow O_2 + NO_2$				-1
R16	$N_2O_5 + h\nu \rightarrow NO_2 + NO_3 + O$	+1	R193	$O + OClO \rightarrow O_2 + O_2$				-1
R24	$ClO + h\nu \rightarrow Cl + O$	+1	R194	$O + OClO + M \rightarrow OCl_3 + M$				-1
R25	$ClOO + h\nu \rightarrow ClO + O$	+1	R195	$O + Cl_2O \rightarrow 2ClO$				-1
R26	$OClO + h\nu \rightarrow ClO + O$	+1	R196	$O + ClNO_3 \rightarrow O_2 + ClONO$				-1
R36	$ClNO_3 + h\nu \rightarrow O + ClONO$	+1	R197	$O + HCl \rightarrow OH + Cl$				-1
R83	$CO_2 + h\nu \rightarrow CO + O$	+1	R198	$O + HOCl \rightarrow OH + ClO$				-1
R95	$2O + M \rightarrow O_2 + M$	-2	R199	$Cl + ClO \rightarrow Cl + ClO$				-1
R99	$O + O_3 \rightarrow 2O_2$	-2	R200	$O(^1D) + HCl \rightarrow Cl + OH$				-1
R102	$O(^1D) + O_3 \rightarrow O + O_2$	-2	R201	$O(^1D) + HCl \rightarrow ClO + H$				-1
R104	$O + H_2 \rightarrow OH + H$	-1	R202	$CFCl_2 + OH \rightarrow CFCl_2 + OH$				-1
R105	$O + OH \rightarrow O_2 + H$	-1	R203	$O(^1D) + CHF_2Cl \rightarrow CF_2Cl + OH$				-1
R106	$O + HO_2 \rightarrow OH + O_2$	-1	R204	$O(^1D) + CCl_4 \rightarrow CCl_3 + ClO$				-1
R107	$O + H_2O_2 \rightarrow OH + HO_2$	-1	R205	$O(^1D) + CFCl_3 \rightarrow CFCl_2 + ClO$				-1
R108	$O(^1D) + H_2 \rightarrow H + OH$	-1	R206	$O(^1D) + CF_2Cl_2 \rightarrow CF_2Cl + ClO$				-1
R109	$O(^1D) + H_2O \rightarrow 2OH$	-1	R207	$O(^1D) + CF_3Cl \rightarrow CF_3 + ClO$				-1
R112	$H + O_3 \rightarrow OH + O_2$	-1	R217	$O(^1D) + CF_3CFCl_2 \rightarrow Cl + 4F + ClO$				-1
R115	$H + HO_2 \rightarrow H_2O + O$	+1	R219	$O(^1D) + COCl_2 \rightarrow 2ClO + CO$				-1
R116	$2OH \rightarrow H_2O + O$	-1	R220	$O(^1D) + COFCl \rightarrow ClO + FCO$				-1
R118	$OH + O_3 \rightarrow OH_2 + O_2$	-1	R220	$Cl + O_2 \rightarrow Cl + O_2$				-1
R122	$HO_2 + O_3 \rightarrow NO + O_2$	-1	R259	$ClO + O_3 \rightarrow Cl + O_2$				-1
R125	$O + NO_2 + M \rightarrow NO_3 + M$	-1	R313	$ClO + O_3 \rightarrow Cl + CO_2$				-1
R126	$O + HNO_3 \rightarrow OH + NO_3$	-1	R315	$O + ClCO_3 \rightarrow Cl + O_2 + CO_2$				-1
R127	$O + HO_2NO_2 \rightarrow OH + NO_2 + O_2$	-1	R316	$O + Cl_4 \rightarrow Cl_3 + OH$				-1
R128	$O + N_2O_5 \rightarrow 2NO_2 + O_2$	-1	R331	$O + H_2CO \rightarrow OH + HCO$				-1
R129	$N_2O_5 + M \rightarrow N_2O + M$	-1	R353	$O(^1D) + CH_4 \rightarrow CH_3 + OH$				-1
R130	$O(^1D) + N_2 + M \rightarrow 2NO$	-1	R358	$O(^1D) + CH_4 \rightarrow CH_3 + H$				-1
R131	$O(^1D) + N_2O \rightarrow N_2 + O_2$	-1	R491	$O(^1D) + CH_4 \rightarrow H_2CO + H_2$				-1
R132	$O(^1D) + N_2O \rightarrow CH_3ONO + O_2$	-1	R496	$O(^1D) + CH_4 \rightarrow H_2CO + O_2$				-1
R133	$O_3 + CH_3ONO \rightarrow NO + O_2$	+1	R497	$O(^1D) + CH_4 \rightarrow CH_3 + O_3$				-1
R143	$N + O_3 \rightarrow NO + O_2$	-1	R498	$O(^1D) + CH_4 \rightarrow CH_3 + O_3$				-1
R144	$N + O_3 \rightarrow NO + O_2$	-1	R532	$O(^1D) + CH_4 \rightarrow CH_3 + O_3$				-1
			R535	$O(^1D) + CH_4 \rightarrow CH_3 + O_3$				-1

Figure 18. Model daily maximum ClO mixing ratios using standard chemistry (solid line) and modified chlorine chemistry (dashed line) for a 960 K parcel. The symbols represent MLS zonal mean ClO interpolated to the latitude and pressure of the parcel, along with the estimated 15% measurement uncertainty. Refer to Figure 17c for the latitude and pressure history of the parcel. MLS zonal means northward of 50° are not shown due to lack of zonal symmetry.

Figure 19. Average model (circles) vs. observed (triangles) daily changes in ozone over the period of the trajectories. Points are slightly offset from the 740, 840, and 960 K potential temperature levels for ease of visualization. The error bars extend out to one sample standard deviation in each direction.

Table 2. Species in Photochemical Model

O	O(¹ D)	H	H ₂	OH	HO ₂
H ₂ O ₂	N	NO	NO ₂	NO ₃	N ₂ O ₃
N ₂ O ₄	N ₂ O ₅	HNO	HNO ₂	HO ₂ NO ₂	CH ₃ ONO
CH ₃ ONO ₂	CH ₃ O ₂ NO ₂	Cl	Cl ₂	ClO	ClOO
OClO	ClO ₃	Cl ₂ O	Cl ₂ O ₂	Cl ₂ O ₃	ClNO
ClNO ₂	ClONO	ClNO ₃	HCl	HOCl	CH ₂ Cl
CCl ₃	CFCl ₂	CF ₂ Cl	CH ₃ Cl	CH ₂ FCl	CHFCl ₂
CHF ₂ Cl	CCl ₄	CFCl ₃	CF ₂ Cl ₂	CH ₃ CCl ₃	CF ₃ CCl ₃
CF ₃ CFCl ₂	ClCO	CHClO	COCl ₂	COFCl	CH ₂ ClO
CH ₃ OCl	CH ₂ ClO ₂	CCl ₃ O ₂	CFCl ₂ O ₂	CH ₂ ClO ₂ H	O ₂ ClNO ₃
CCl ₃ NO ₄	CFCl ₂ NO ₄	CH ₃	CO	CO ₂	HCO
H ₂ CO	CH ₂ OH	CH ₃ O	CH ₂ O ₂	CH ₂ OOH	CH ₃ OH
CH ₃ O ₂	CH ₃ OOH	HOCH ₂ OO			

Figure 1.

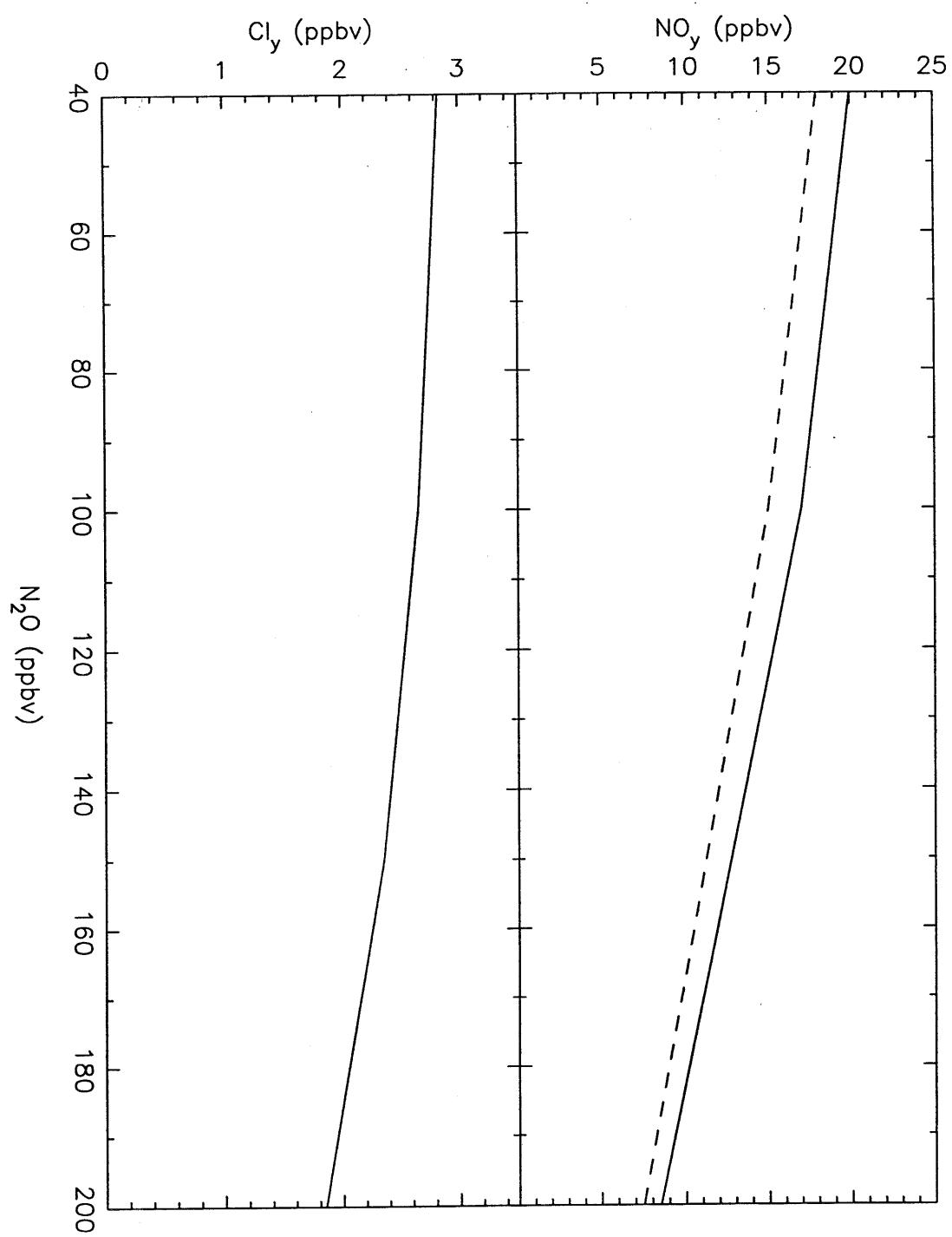
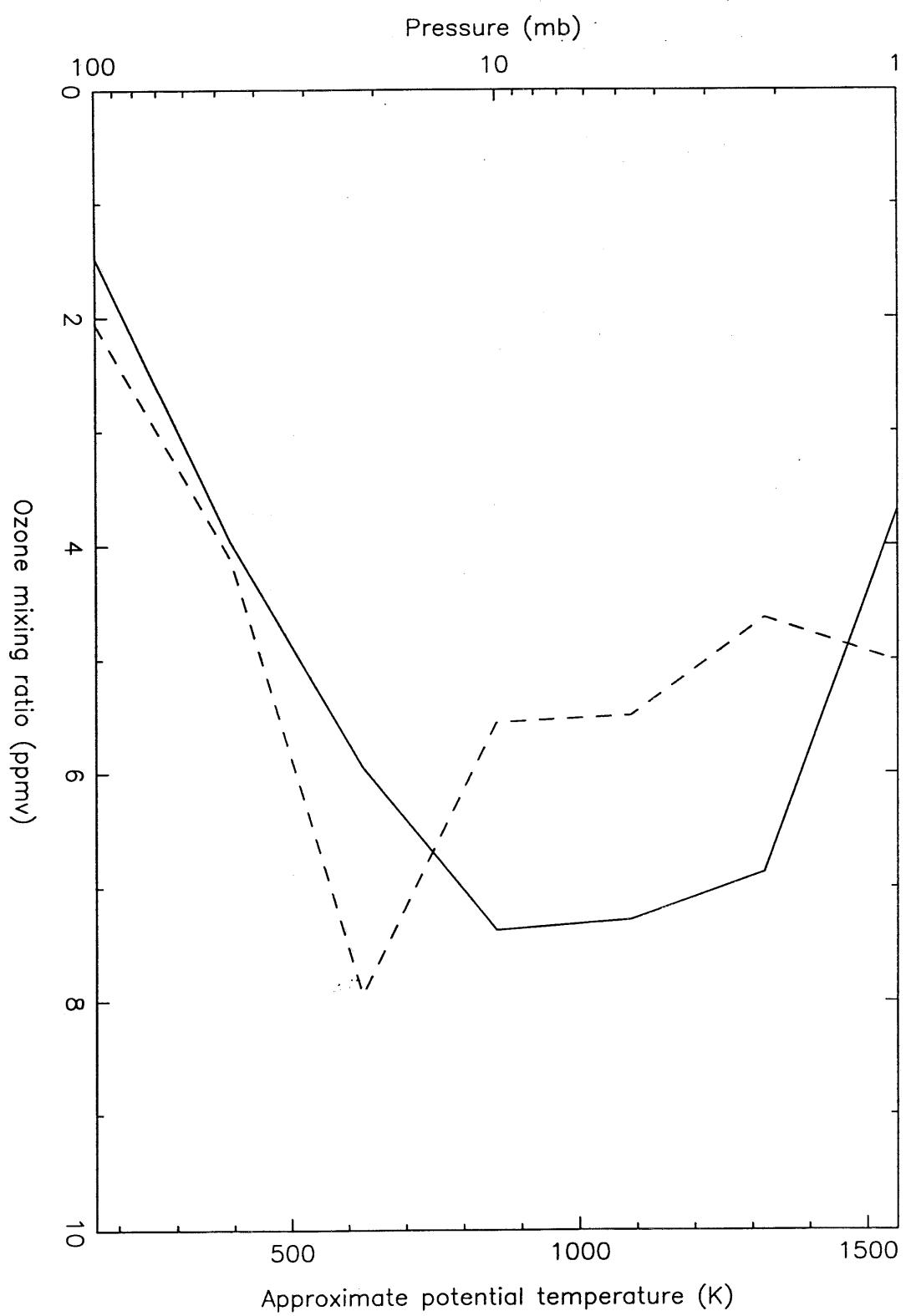


Figure 2.



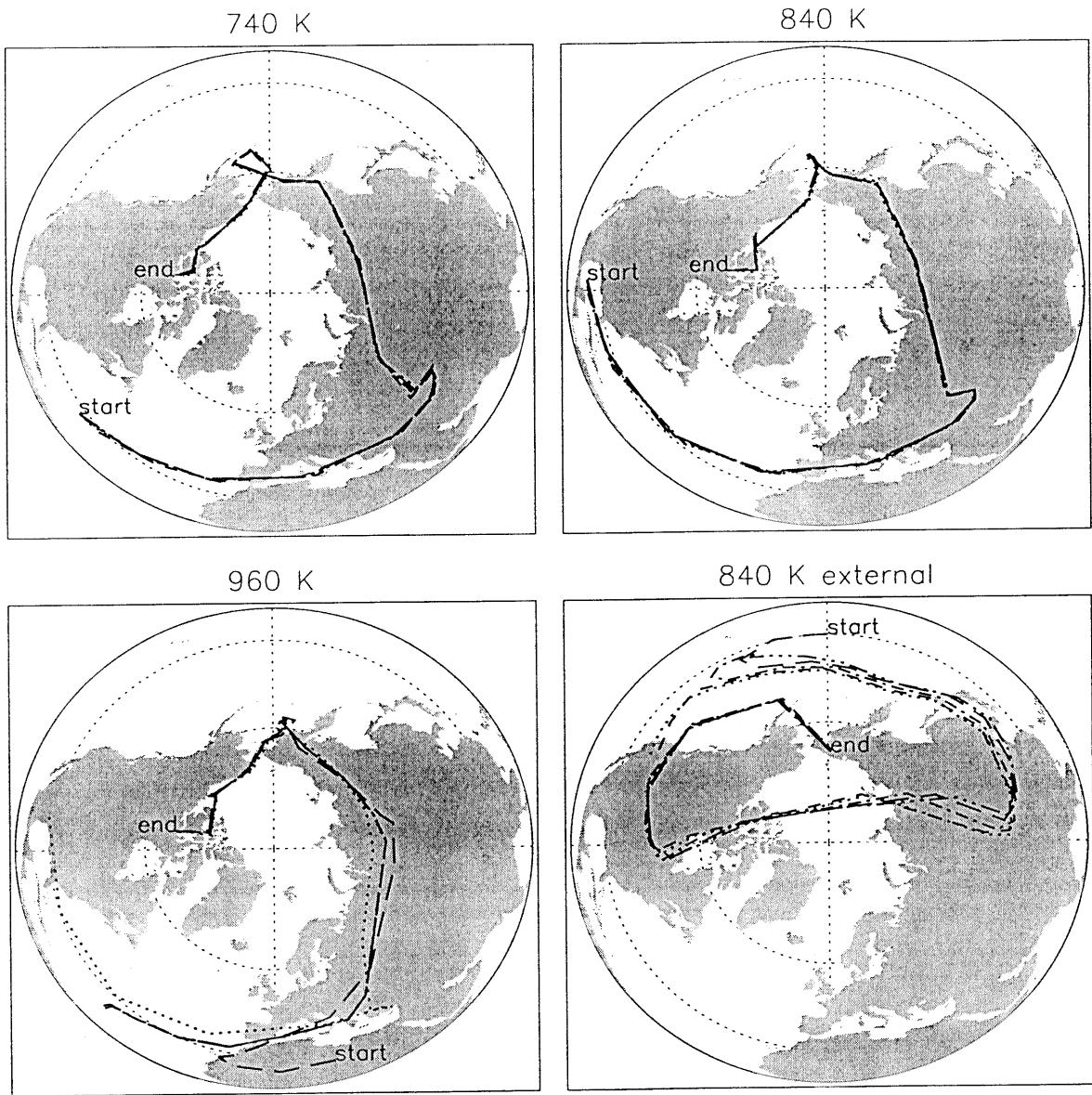


Figure 3.

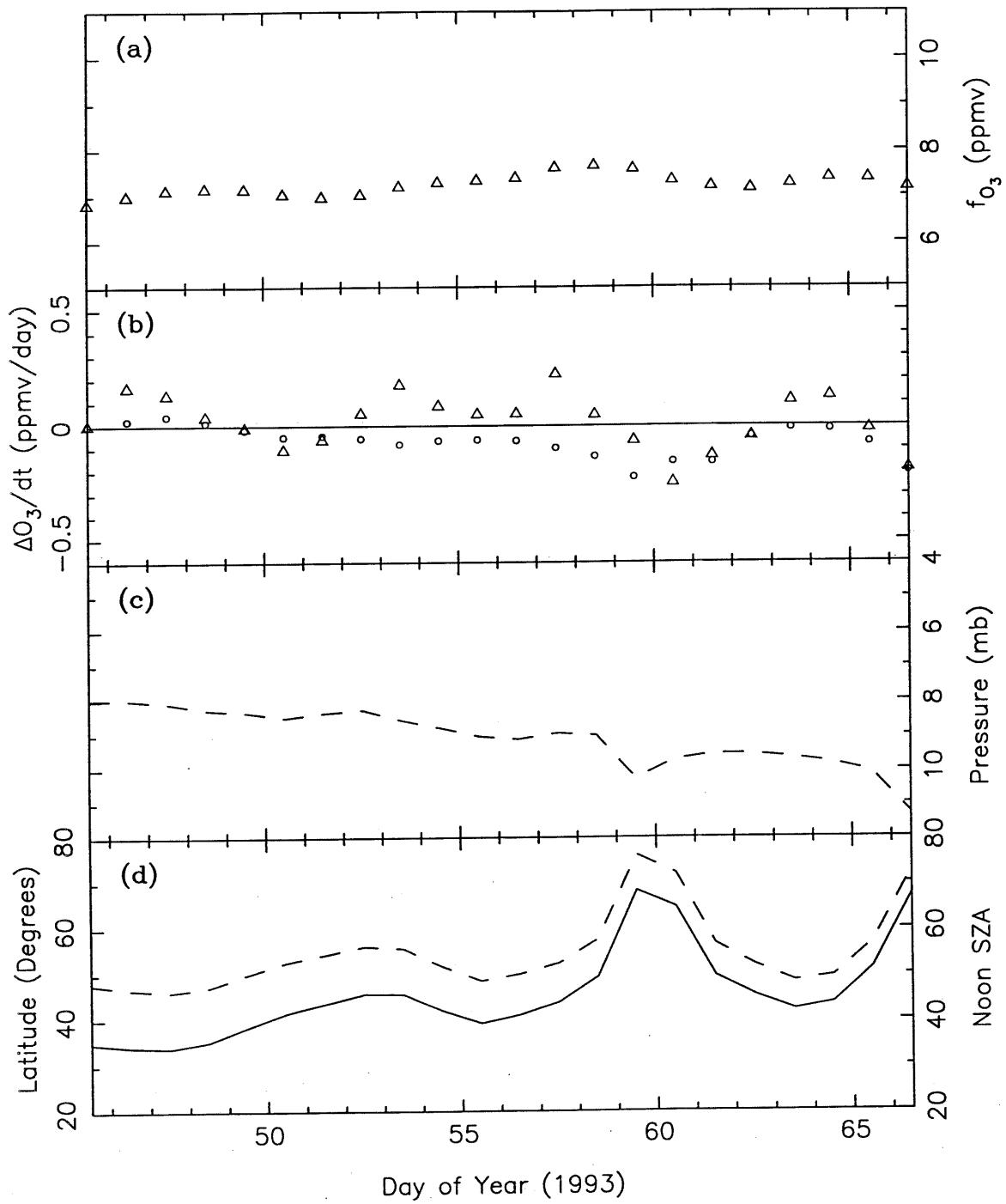
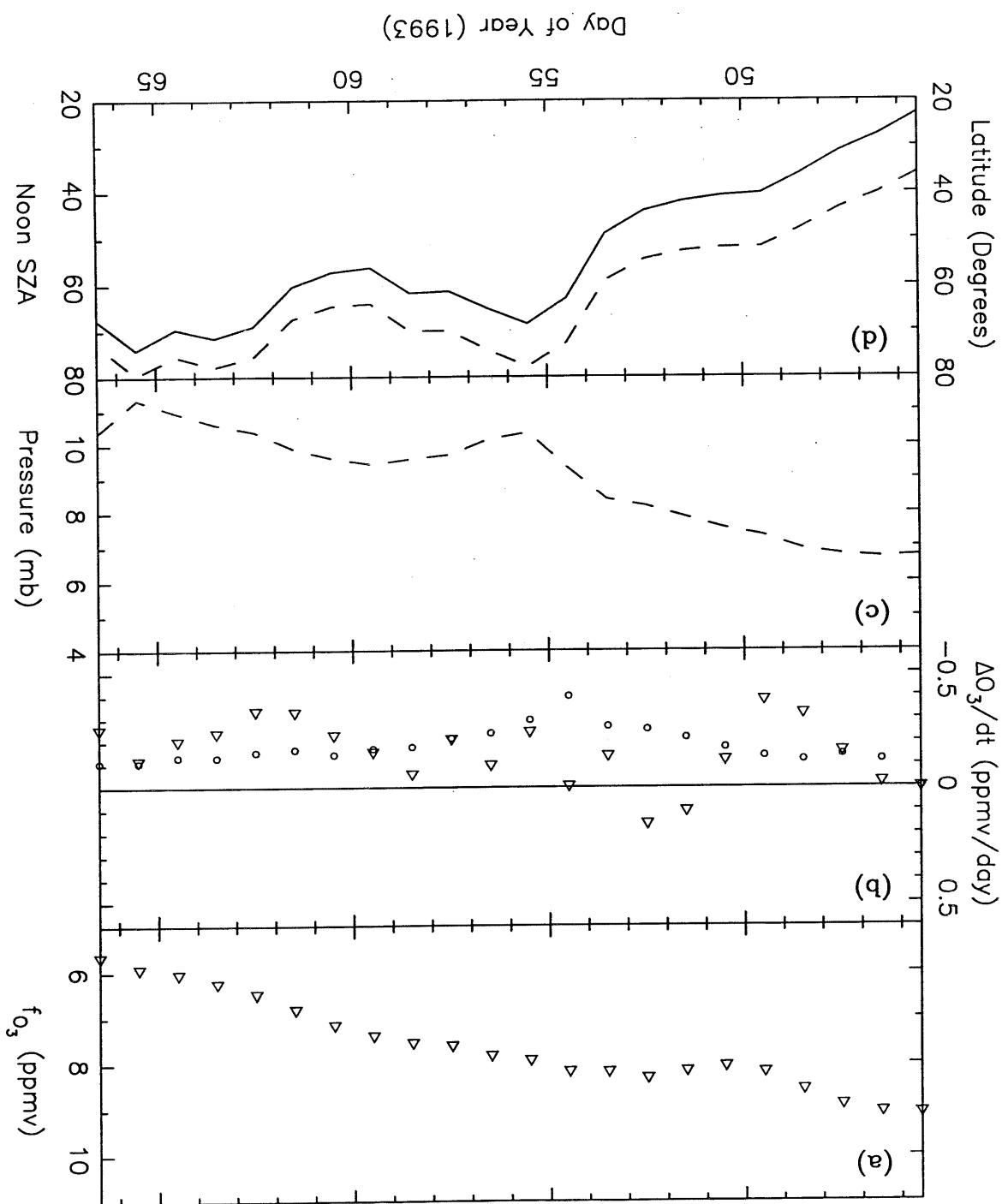


Figure 5.

Figure 4.



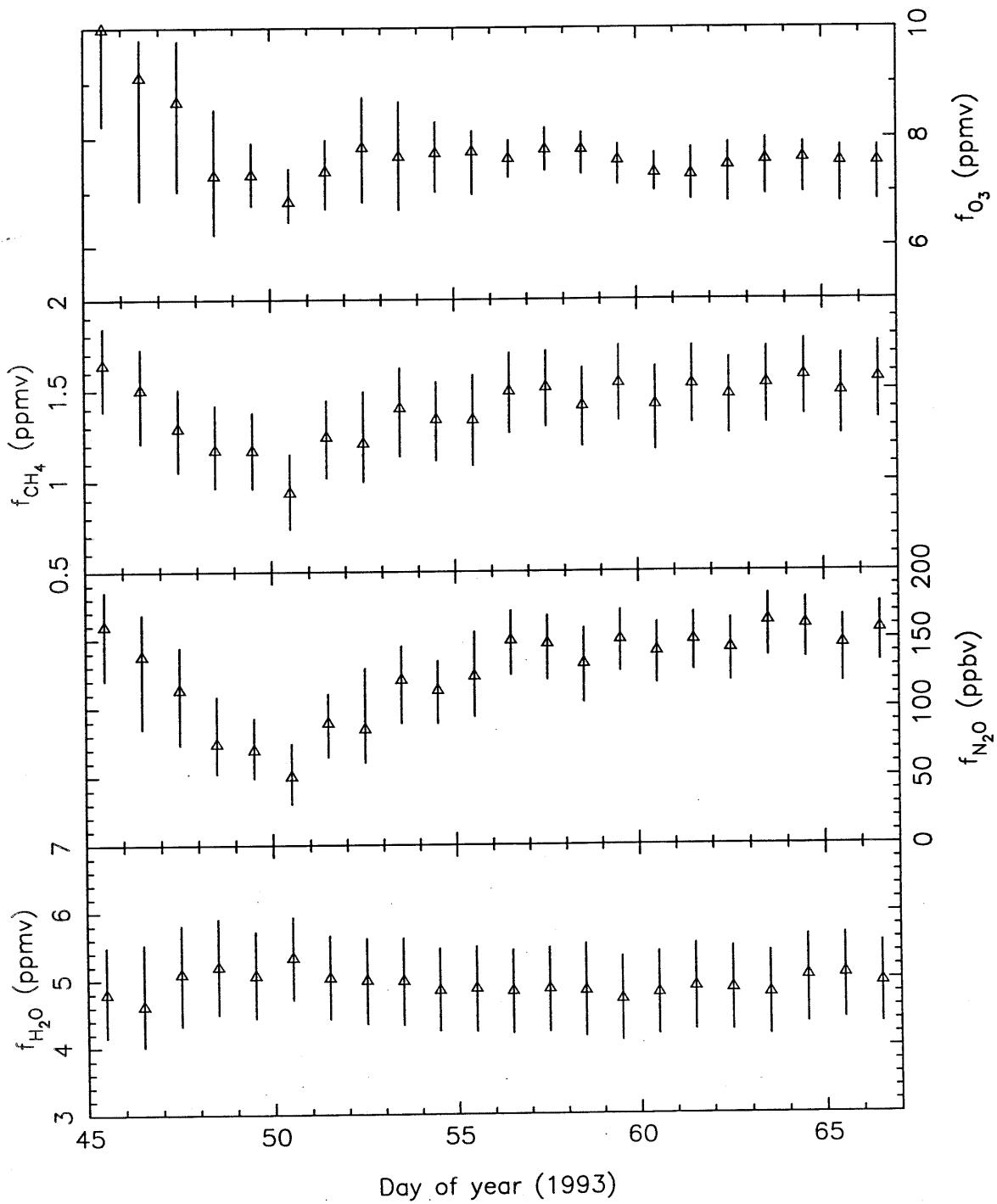


Figure 7.

Figure 6.

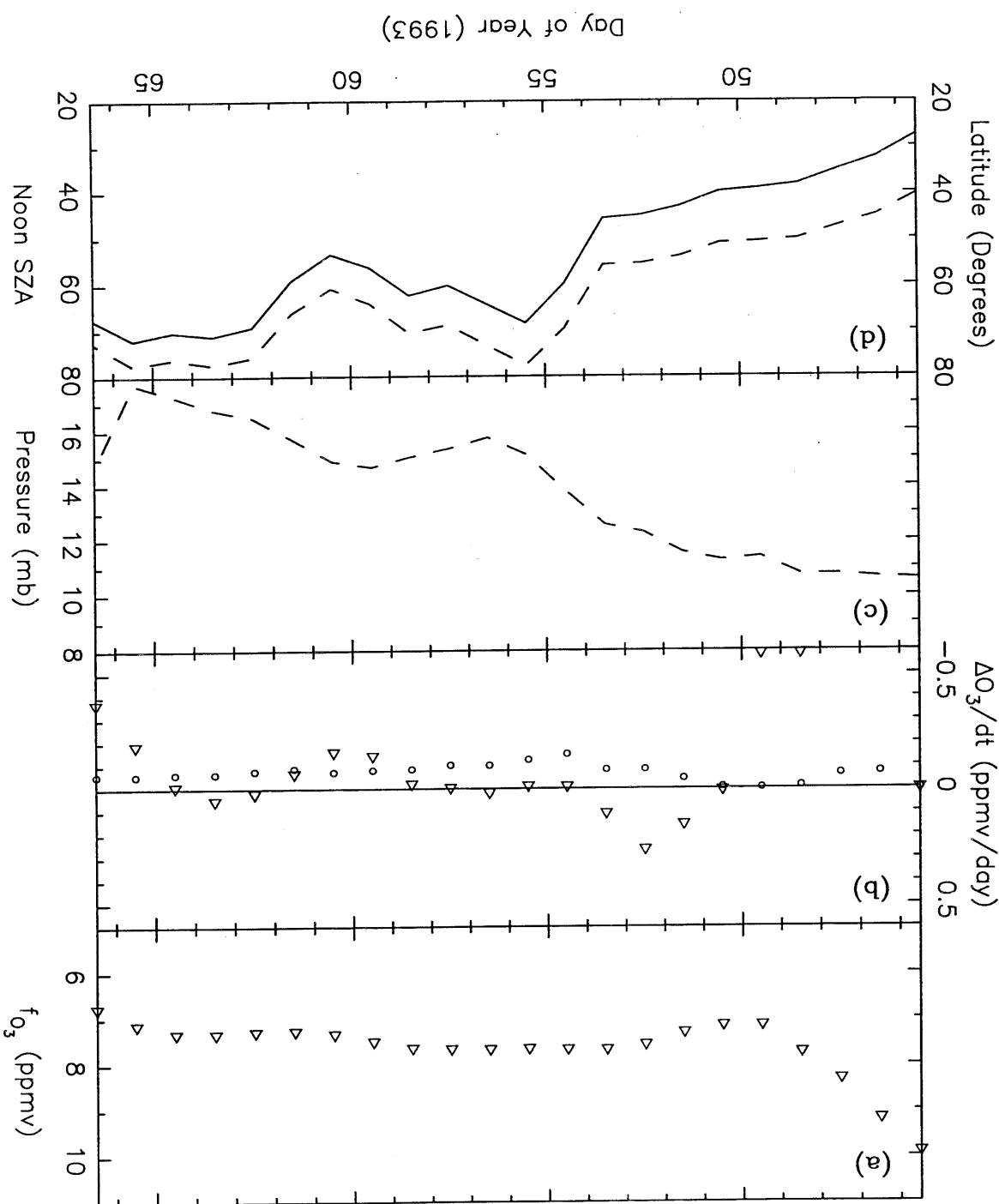
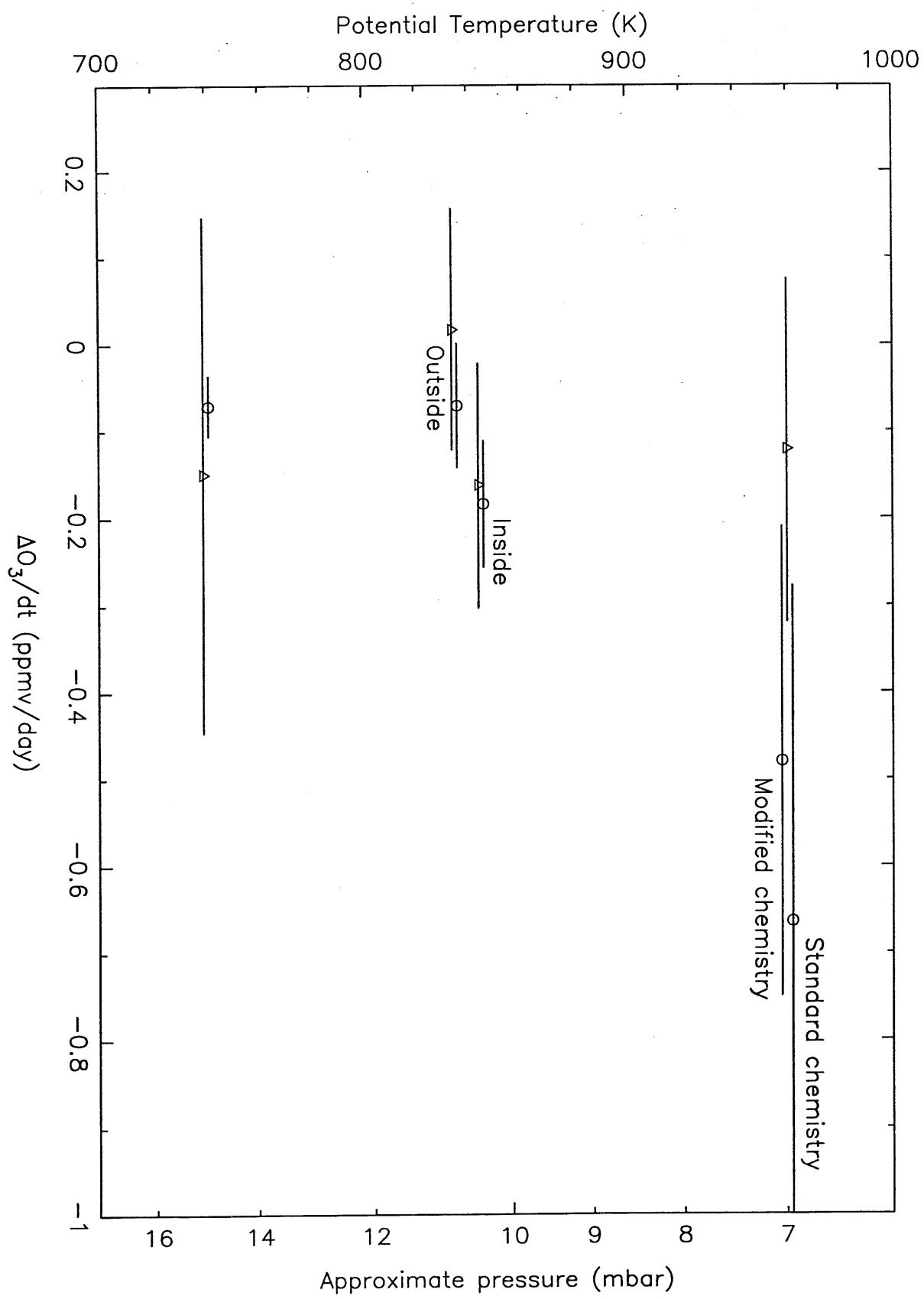


Figure 9.



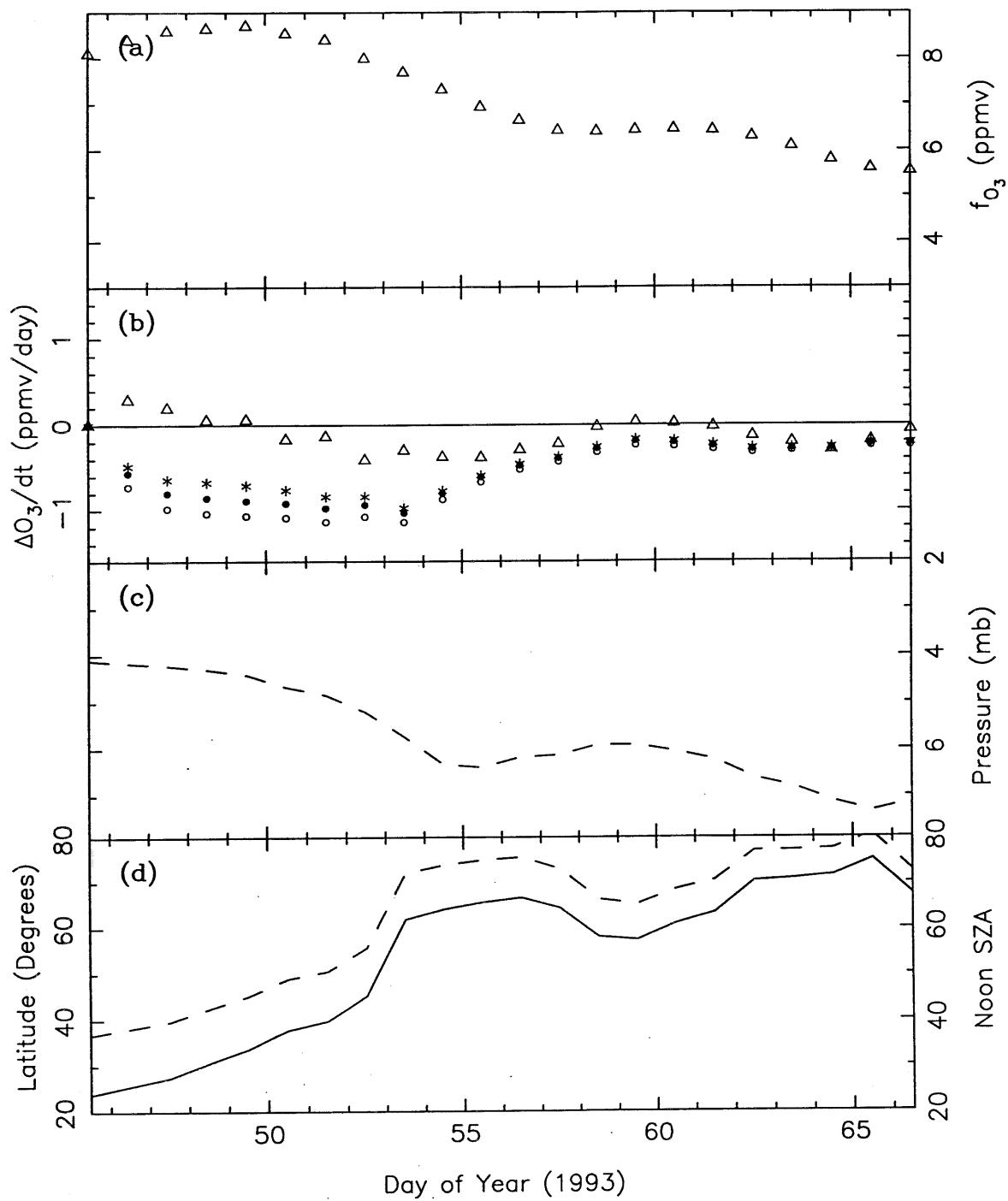


Figure 8.

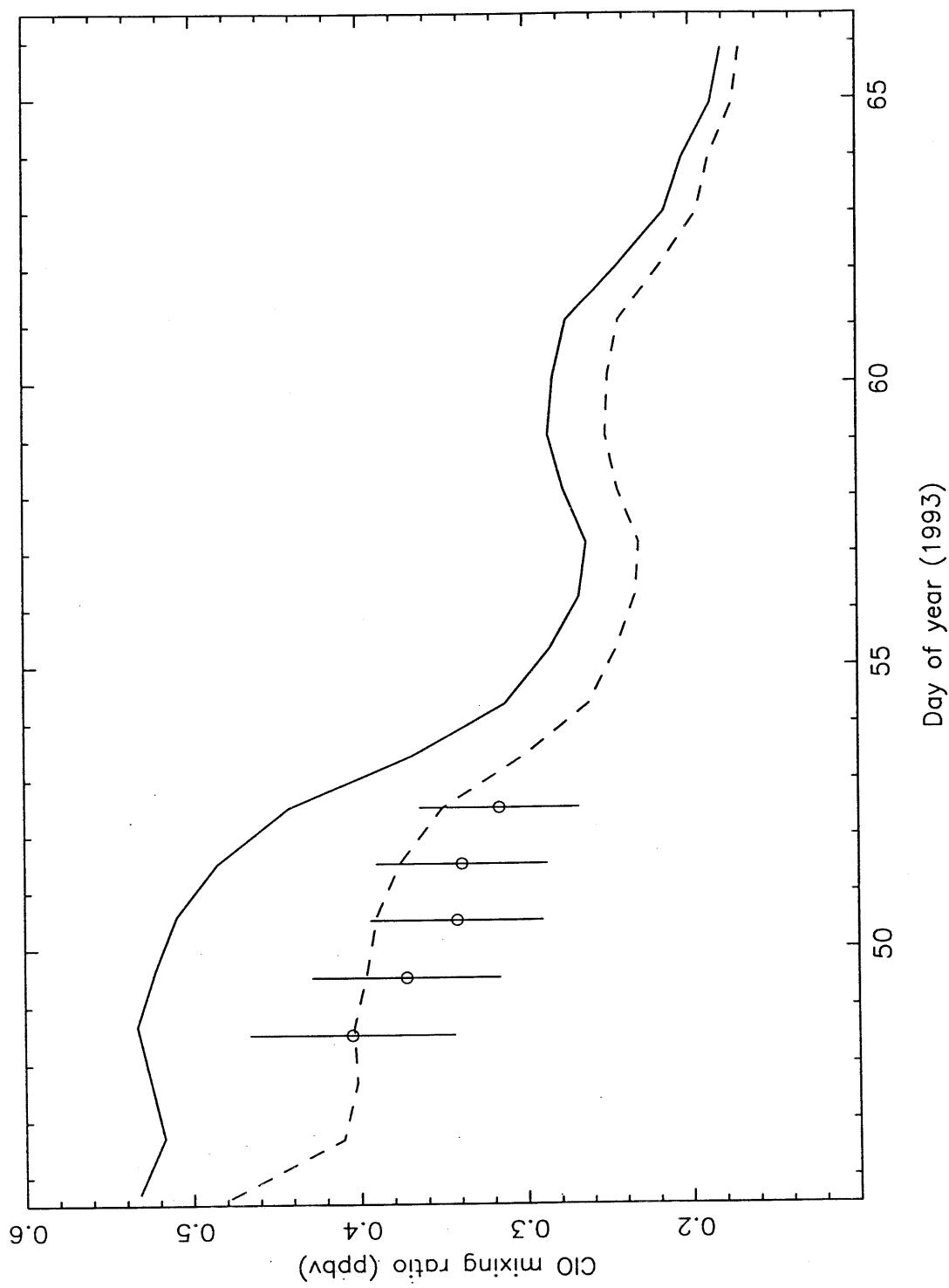
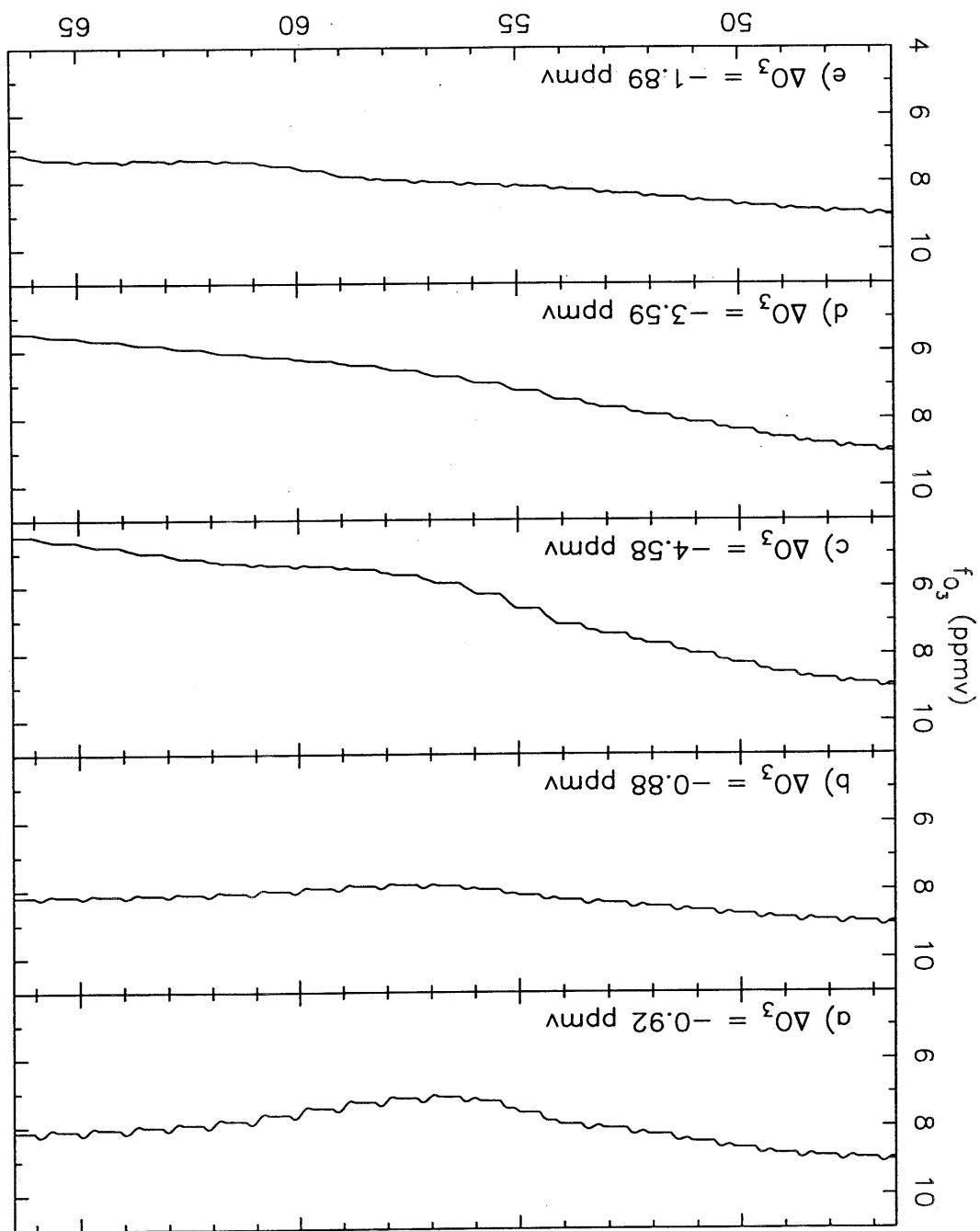


Figure 11.

Figure 10.

Day of Year (1993)



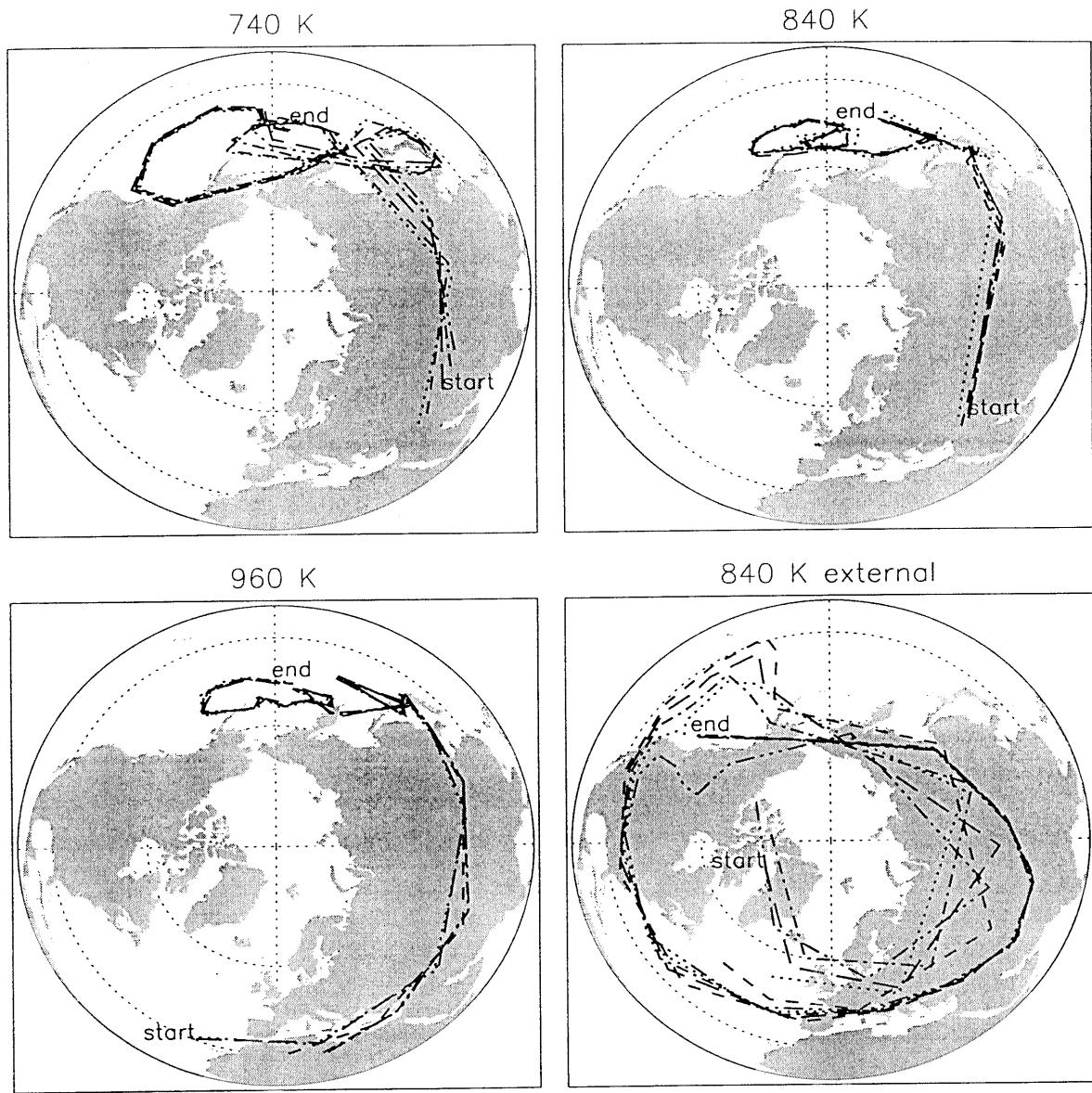
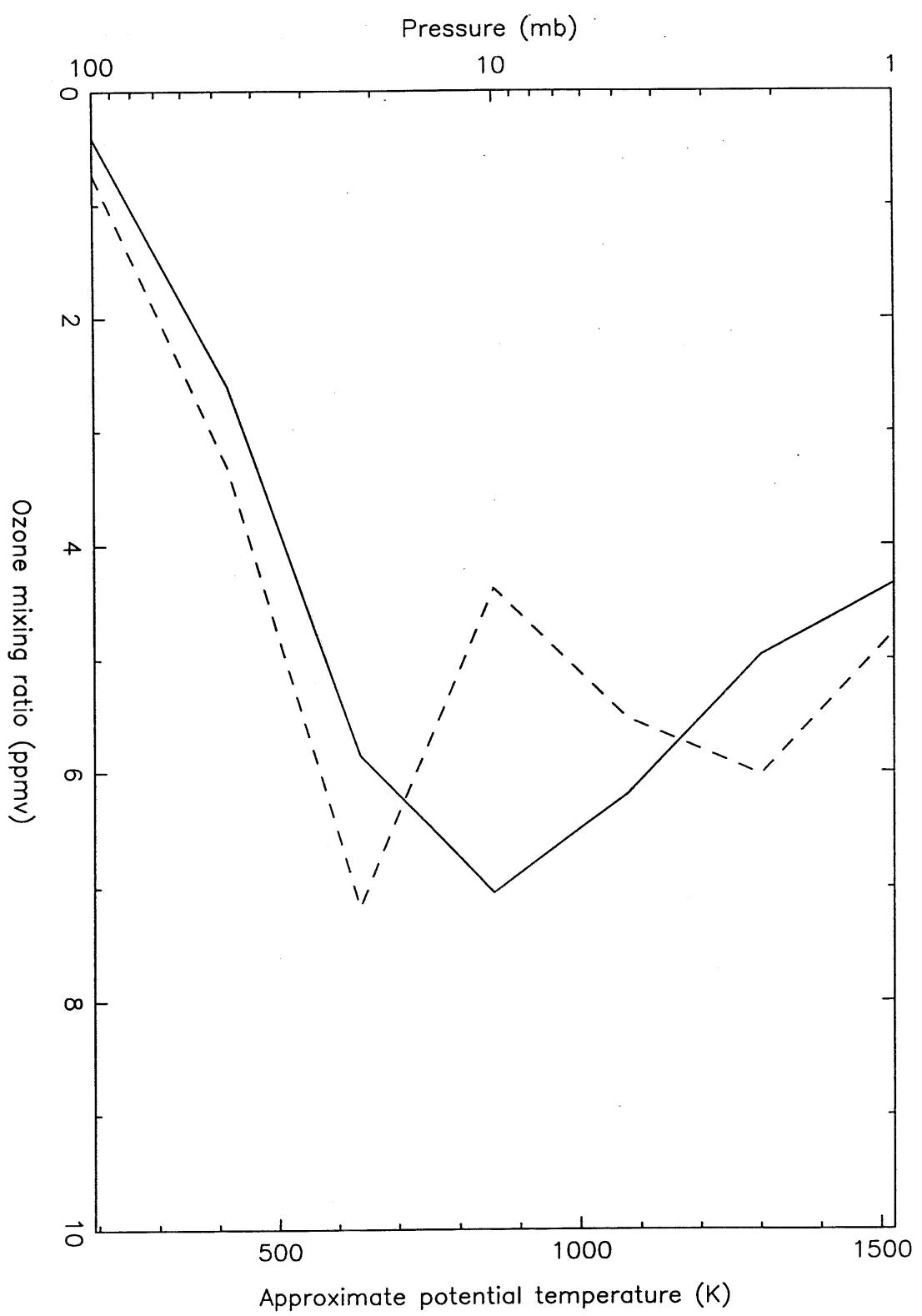


Figure 13.

Figure 12.



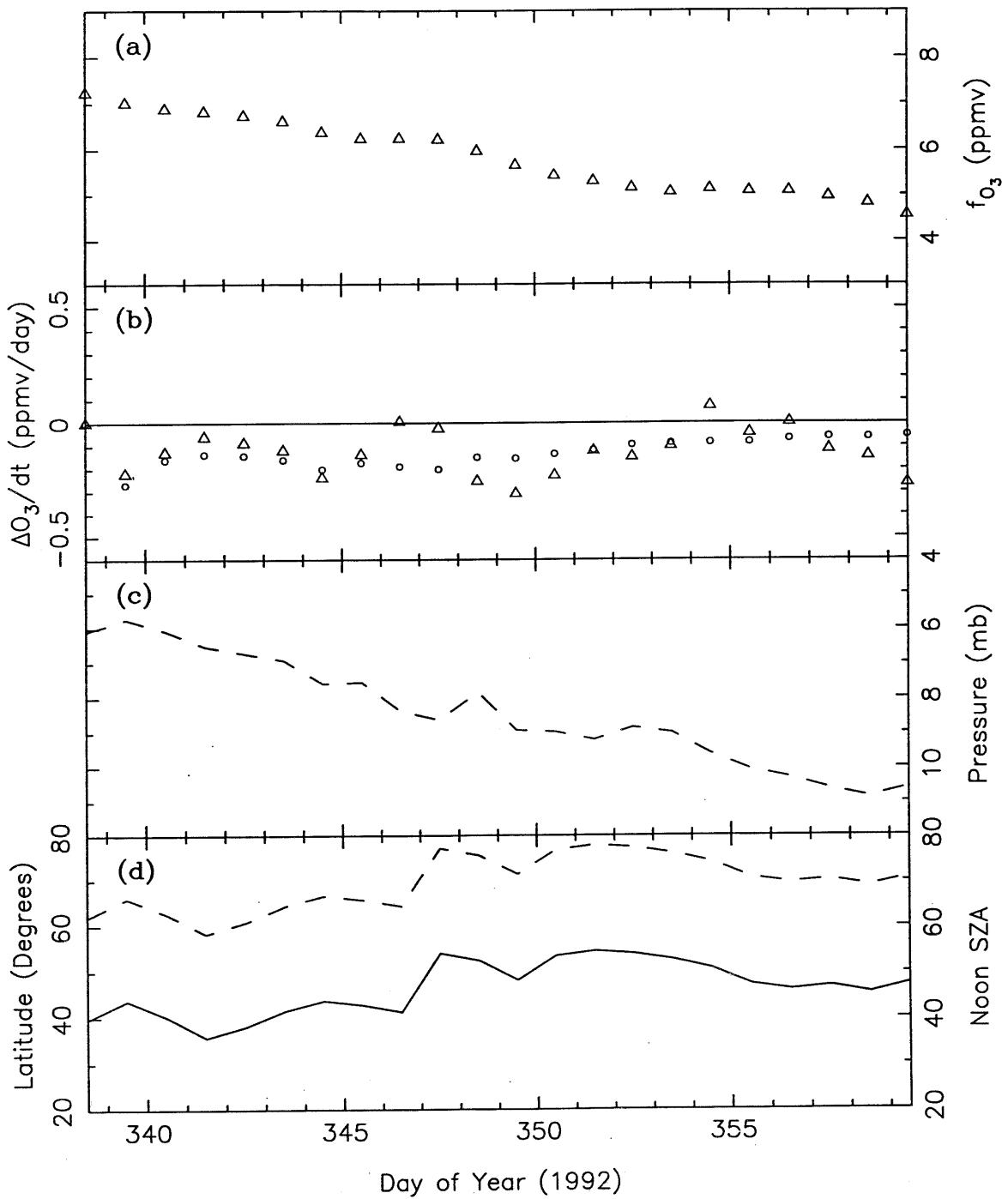
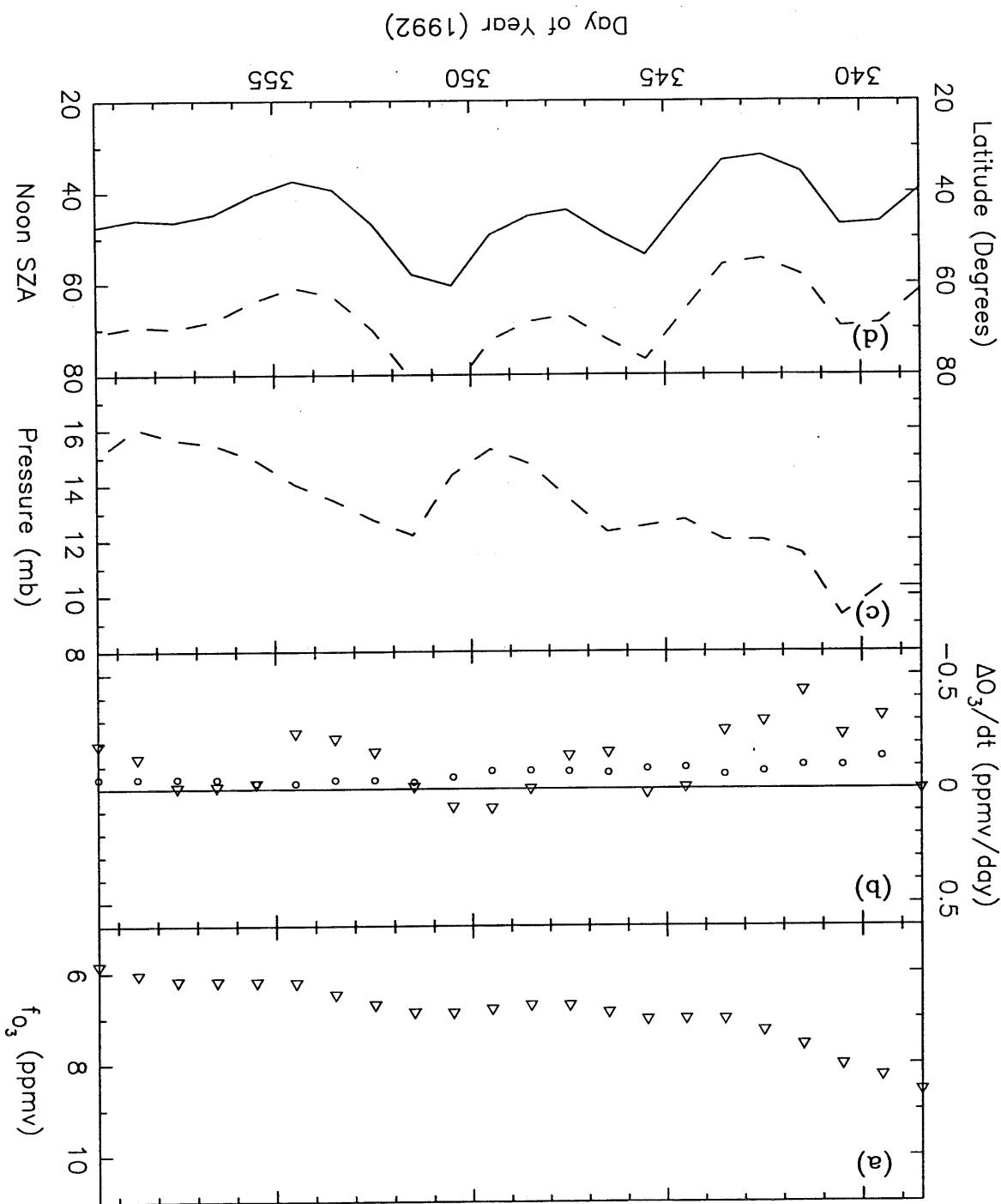


Figure 15.

Figure 14.



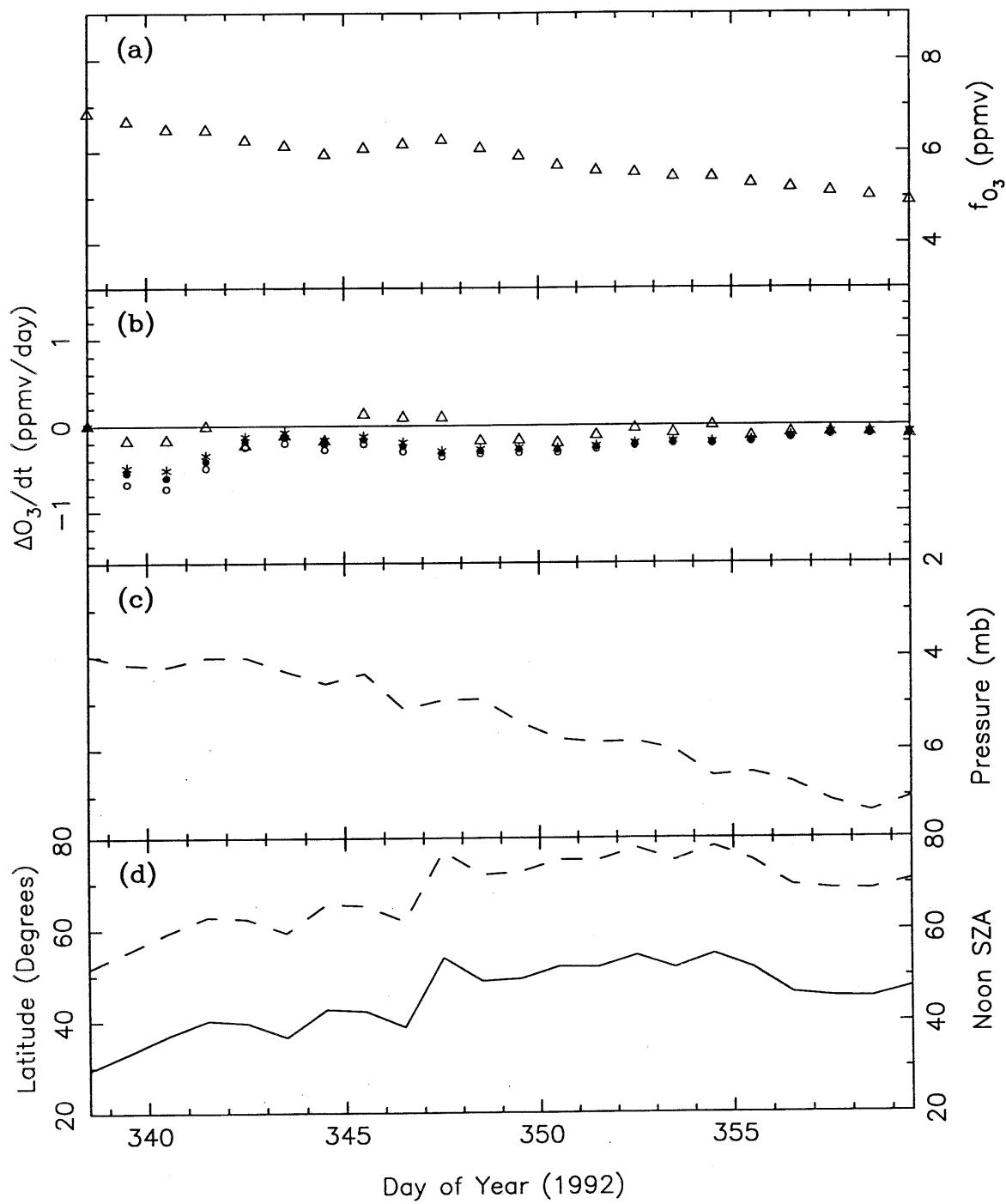


Figure 17.

Figure 16.

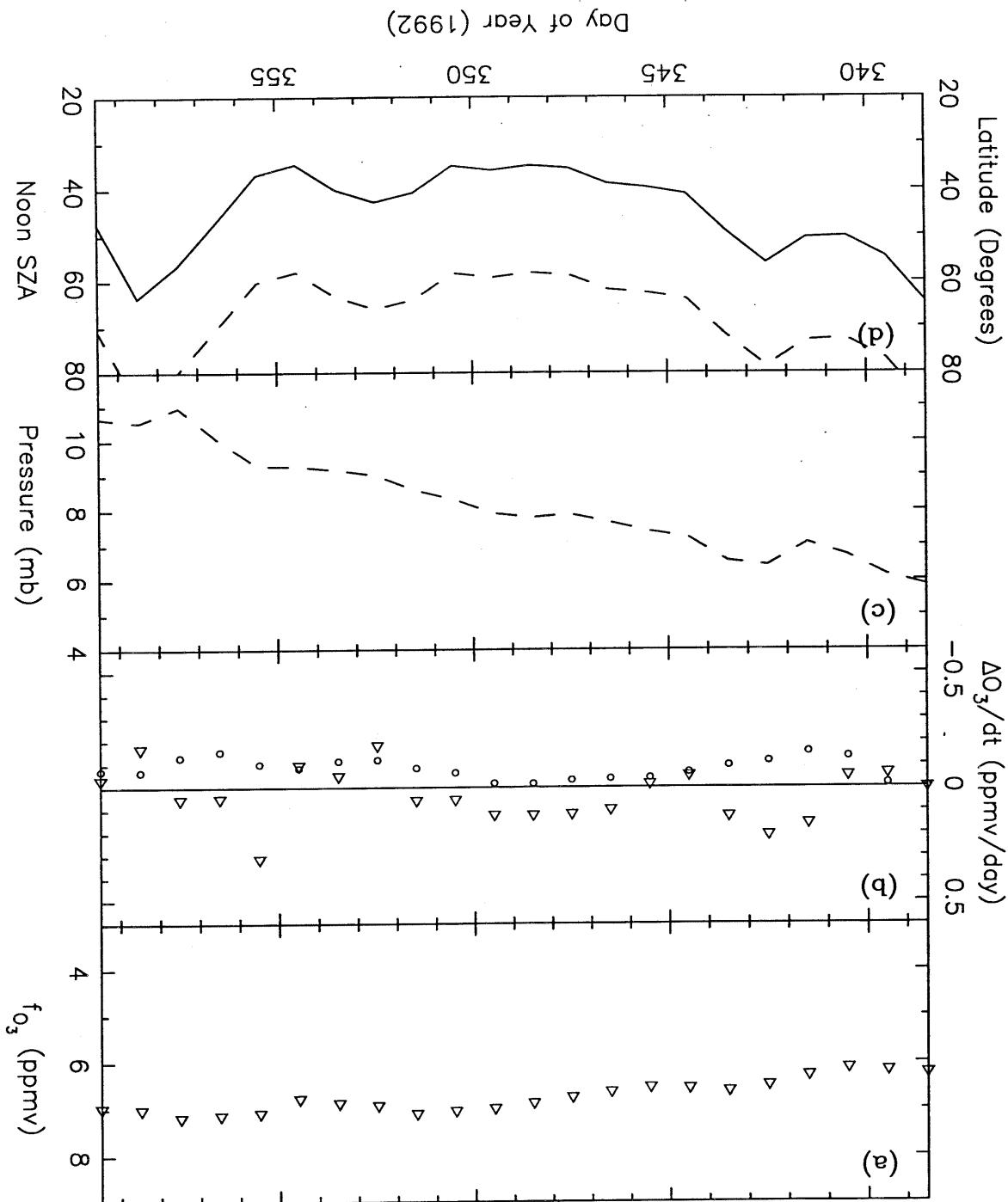
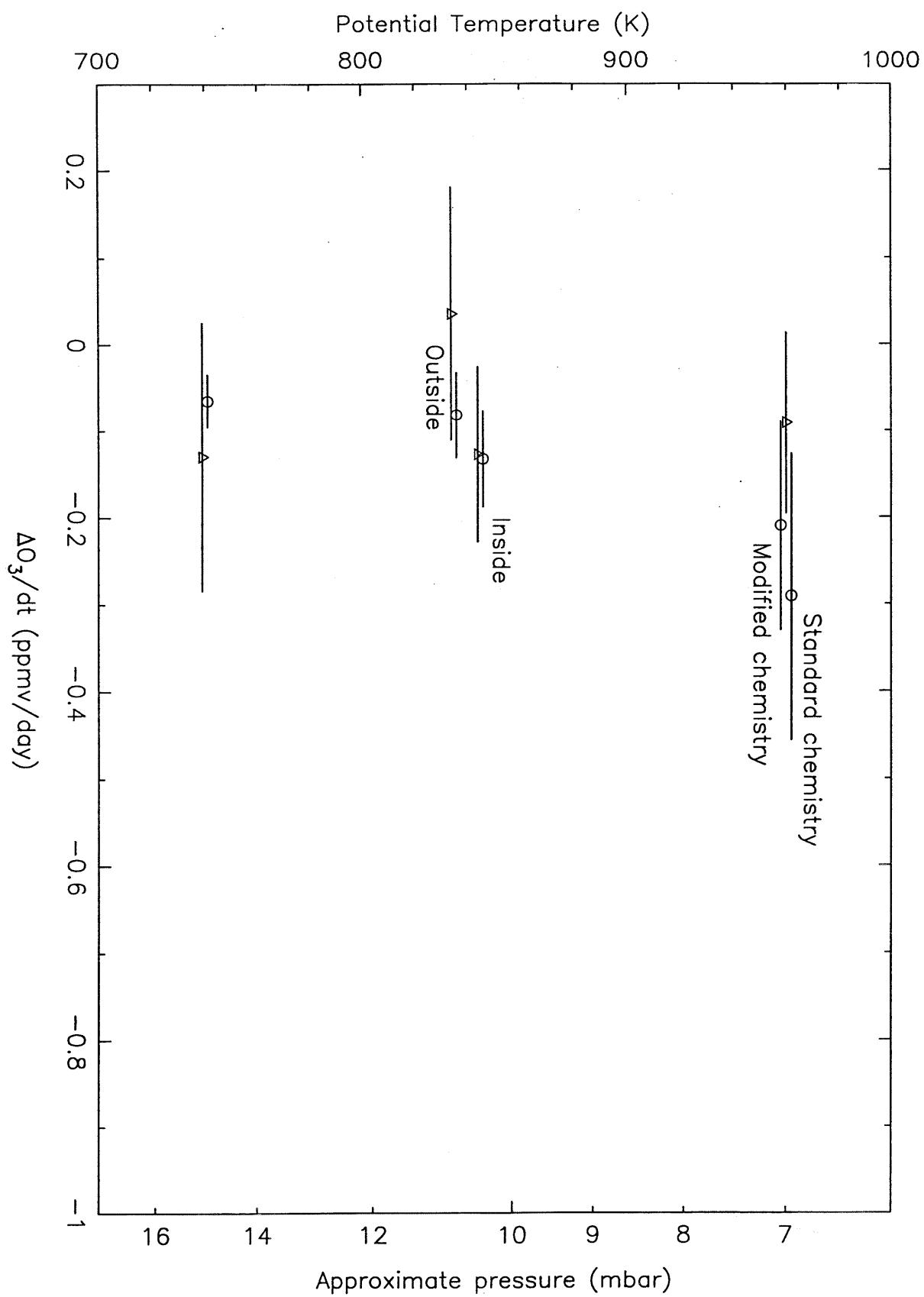


Figure 19.



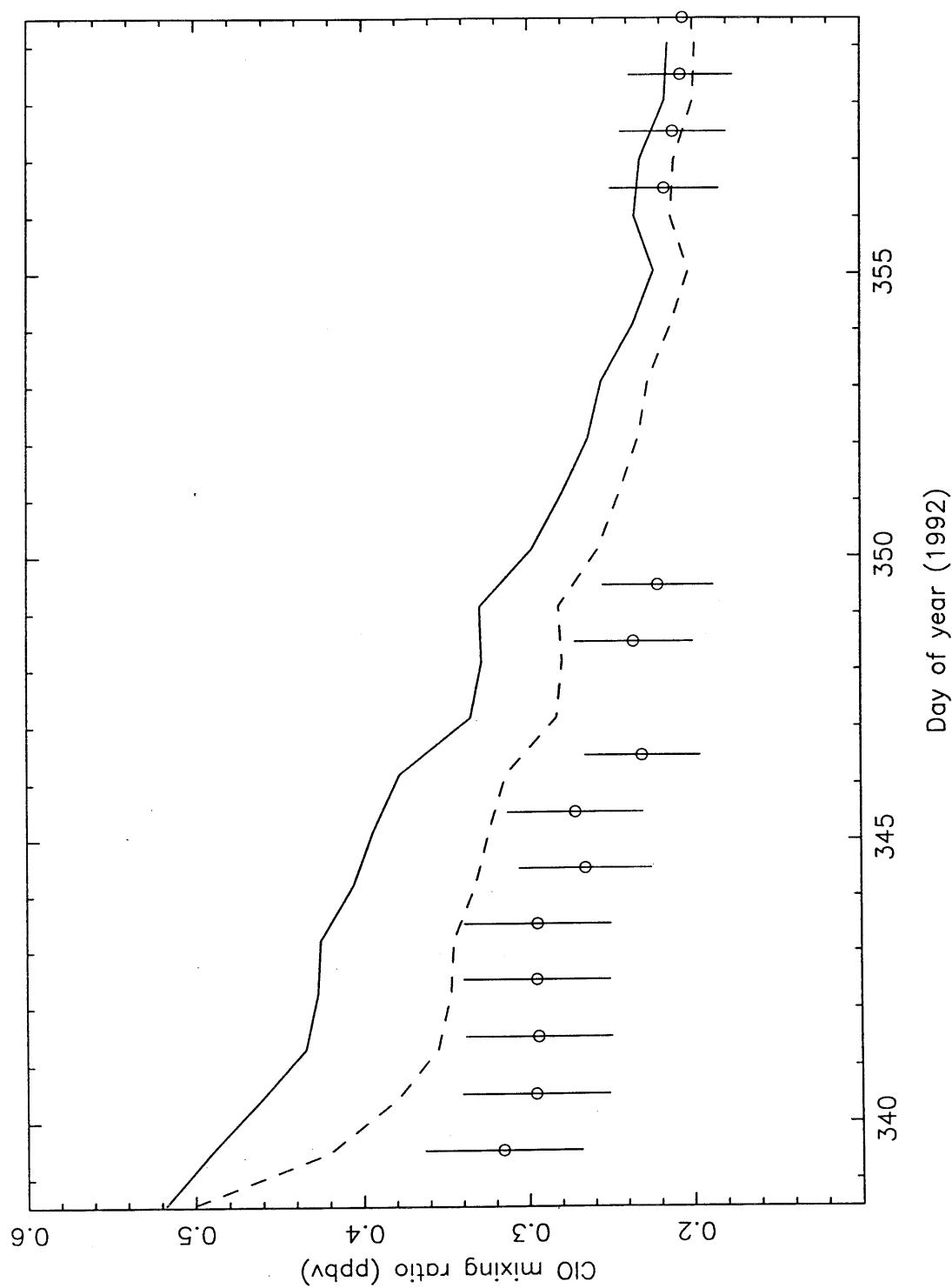


Figure 18.